HEAVENS SURVEY'D,

ANDTHE

True System of the Universe

Delineated, so as to form a Curious

ASTRONOMICAL INSTRUMENT:

EXHIBITING

The Number, Order, Periods, Aphelions, Perihelions, Nodes, &c. of all the PRIMARY PLANETS by Inspection:

TOGETHER WITH

An easy and expeditious Method of ascertaining by the same Instrument, and a new Set of Tables, their Anomalies, Longitudes, Latitudes, Retrogradations, Conjunctions, Elongations, and Distances at all Times, both Heliocentric and Geocentric; that is, as seen from the Sun and from the Earth; and also from one another.

In the Course of the Work is met with

A View of the System from the Earth, elucidating and explaining the true Causes of the Directions and Retrogradations of the Planets; and of the various Phases they exhibit to us here: With the Reasons of the Transits of Venus and Mercury over the Sun; the Times when they happen;—and an easy Method of delineating the Tracks those Planets take in passing over his Disk.——Also, the remarkable Path which the Planet Mars seemed to describe in the Heavens in the year 1762.

The whole adapted to the NEW STYLE;

And constructed in a Manner so easy and natural, as to convey to the Astronomical Learner a perfect Knowledge of the Solar System at first View.

Addressed to the YOUTH of GREAT BRITAIN and IRELAND.

By the Reverend Mr. TURNER, late of Magdalen Hall, Oxford;

Author of the View of the Heavens; -----View of the Earth; -----Trigonometry rendered Easy and Familiar; ------System of Gauging; -----Chronologer Perpetual; ----- and a New Introduction to Book-Keeping.

Rector of Comberton; Vicar of Elmley; Minister of Norton, and Chaplain to the Right.
Honourable the Countess Dowager of Wigton.

Cælique Meatres

Describent Radio; et surgentia Sidera dicent.

VIRG.

LONDON;

Printed for S. CROWDER, at No. 12, Pater-noster-Row; MDCCLXXXIII.

A

VIEW of the HEAVENS:

Being a Short, but Comprehensive,

SYSTEM

OF

MODERN ASTRONOMY.

EXHIBITING,

- I. The Number, Order, Distances, Magnitudes, and Periods of all the Planets and their several Moons, composing our System, which the Learning of the present Age esteems as so many Worlds full of Inhabitants.
- II. The length of the Day and Year, with the Variety of the Seasons in each Planet; and alfo the Phænomena of the Heavens to the Inhabitants thereof.
- III. Some Account of the Comets, their Number, Period, and Appearances; and also the Directions of their fiery Trains thro' the Heavens; with probable Conjectures of the Uses of those amazing Bodies.
- IV. The Number, Magnitude, and Distances of the Fixed Stars; with their Divisions into Signs, Catalogues, and Constellations.
- V. The true Figure of the Planets' Orbits, with the Nature of the Motions in them; their

- Aphelions, Perihelions, Eccentricities, Nodes, &c. &c.
- VI. The Inequality in the Length of the Natural Day, commonly called the Equation of Time.
- VII. The Power that retains the Earth and Planets in their Orbits.
- VIII. The direct and retrograde Motions of all the Planets; as also, the Reasons why they sometimes appear stationary, or not to move at all.
- IX. The Nature and Causes of Eclipses, both of the Sun and Moon; with an easy and expeditions Method of calculating the Eclipses which will happen in any Year.
- X. The Description and Use of a curious Astronomical Clock, which will shew the Flour of the Night by the Stars.

TO WHICH IS ADDED,

The Use of the CÆLESTIAL GLOBE:

WITH

Its Application to a Number of very interesting Problems. Concluding with some curious Phænomena upon the Sun and Moon exhibited in a darkened Room; and a sew select Paradoxes, intended to excite the Attention of the Learner.

The whole illustrated with Copper-plates of the System, the Sun, Moon, Eclipses, &c. and disposed in to easy and natural a Manner, as to be understood in a few Days.

THE SECOND EDITION, WITH MANY ADDITIONS AND IMPROVEMENTS.

By the Rev. Mr. TURNER, of Magdalen-Hall, Oxford;

Rector of Comberton; —Vicar of Elmley; —Minister of Norton; —and Chaplain to the Right Honourable the Countess Dowager of Wigton. —Author of The Heavens Survey'd; —The View of the Earth; —Plain Trigonometry rendered Easy and Familiar; —System of Gauging; —Chronologer Perpetual; —and a New Introduction to Book-keeping.

The WORLDS were framed by the Word of God.

PAUL.

LONDON:

Printed for S. Crowder, at No. 12, Pater-noster-Row, MDCCLXXXIII.

T O T H E

READER.

STRONOMY in all Ages has been cultivated not only by Men of the greatest Genius, but those of the highest Honours.—Kings as well as Philosophers have studied the Revolutions of the glowing Worlds above, and made them useful to us, by measuring out our Dates and Time here on Earth.

The vast Advantages derived to Mankind from thence, together with the Pleasures resulting from this Heavenly Science, induced me to draw up this new Work for the Youth of these Kingdoms; which, as it is now become a Science so necessary in civil and facred Life, and of which every Person (who would be a Scholar) should have some Knowledge, I have attempted in a Manner as easy and engaging as possibly I could.

With this View I have supposed the young Astronomer transferred from the Earth to the Sun, where he may the better behold the several Worlds wheel round him, and observe their various Distances, Periods, Dimensions, Inclinations of their Axes and Orbits, Aphelions, Perihelions, and Nodes; and have also supposed him to traverse the System, and to have actually taken those several Measures and Dimensions, and noted them down in a Table.

Next, I have directed him how to delineate these Observations, and afterwards to fit up his Projection into a Curious Astronomical Instrument, by which he may easily calculate, at all Times, the Places of all the Primary Planets, and exhibit their Situations as naturally as in the Heavens themselves.

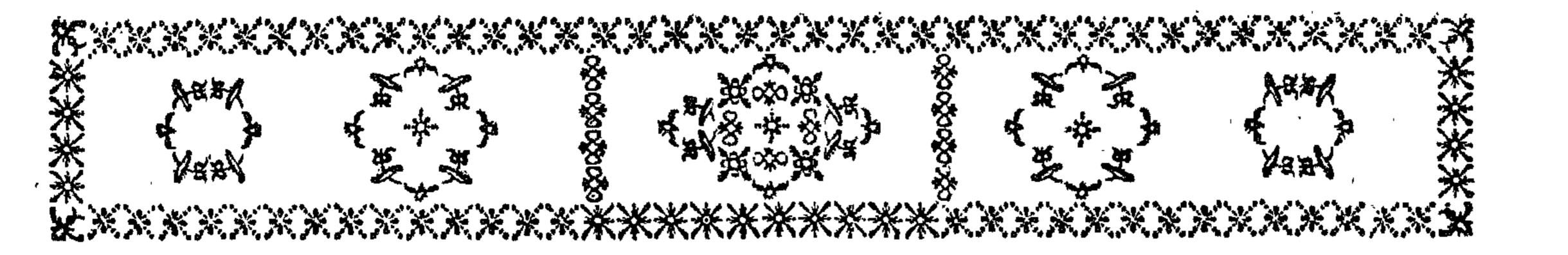
To

To make it still more evident, I have annexed a Variety of Examples—some for each Planet; which, being once perused, he cannot fail of understanding the usual Method of making these Calculations by Triangles, as he will immediately see on the Instrument how the Lines are situate in the Heavens;—what Sides and Angles are given, and what required to be found by Trigonometrical Calculation.

He has in the last Place, a new Sett of Tables for finding the mean Anomalies, and Equation Tables for each Planet, for reducing them to the true Anomalies.—Also, a Table is added of Inclinations, or Heliocentric Latitudes of the Planets, depending on their Distances from their Nodes: And to make the Work more complete—all these Tables are adjusted to the New Style.

That the Youth may receive that Pleasure and Advantage from the Perusal of this Work which was intended them by its Publication, is the sincere Wish of

The AUTHOR.



THE

HEAVENS SURVEYED.

Former of the Universe, than — to contemplate the Planetary Worlds above;—to survey the astonishing Spaces through which they rowl;—to resect on the Vastness of their Bulks;—the amazing Velocity of their Motions;—and above all, the exact Regularity and Harmony of their Periods:—so I flatter myself, that any rational Scheme tendered to the World, with a Design of facilitating Enquiries of such a sublime and interesting Nature, cannot but be pleasing to every thoughtful and ingenuous Mind.

And, though it be very possible to give the Learner a Knowledge of the Wonders and Immensity of that Heavenly Fabric from this Earth on which he now resides; yet as it would be more easy and natural for him to behold it from the Center, round which all the Planets move, he will give me Leave to conduct him thither, and point out to him the wonderful Construction and Phanomena of the System from that Point of View *.

* As the young Astronomer hath pointed out to him here, the Order and Harmony of the System from the Sun, the Center of it; so at Page the 24th, the Diversity of the Motions and Aspects of the several Planets are exhibited to him from the Earth, the Place of his present Residence.

On

On his Arrival there, he will find the Sun, an huge Globe of Light, plac'd in the Midst of immense Space, with Six opaque spherical Bodies moving round him as their Center; which at different Distances, and in different Periods of Time, perform their Revolutions all the same Way, i.e. towards the Left-hand, in the Order following.

First, he will see Mercury the nearest of all the Planets, moving the quickest, and describing the smallest Orbit round the Sun.—Next, he will see Venus, which, as she is surther off, takes a larger Circuit, and a longer Time to compleat it.—Beyond her he will behold our Earth moving in her Orb not quite so fast as Venus.—Next her advances Mars, but revolves much slower.—Then comes, at a vast Distance beyond him, the Planet Jupiter, slower still.—And last of all, he sees Saturn, as it were come creeping round his Orb.*

At a vast Distance beyond Saturn, he will observe the azure Sphere of the fixed Stars, sparkling all over with Lights of different Lustres and Magnitudes. These he will find, (contrary to the Planets) never to vary their Places, but remain always in the same Position. And as it is impossible to distinguish each single Star by a particular Name, Astronomers have agreed to suppose a Number of them lying near each other, to be covered with the Image or Picture of some Animal, or other Thing.—That Round of the starry Arch opposite to the Orbs of the Planets, is divided into 12 Images, whose Names are, Aries—Taurus—Gemini—Cancer—Leo—Virgo—Libra—Scorpio—Sagitarius—Capricornus—Aquarius—Pisces. This Space is called the Zodiac. It is supposed to be divided into 360 equal Parts, called Degrees: Thirty of which belong to each Image.

- The Earth's Orbit passes exactly opposite the Middle of these 12 Signs or Pictures.—And a Plane level with the Earth's Orbit extended from the Sun to the fixed Stars is called the Plane of the Ecliptic : and this Astronomers make the standard to which the Planes of the other Planets (which is very little) are found to incline.
- * Besides these six Primary Planets, which continually circulate round the Sun as the Center of their Motion, there are several Smaller which circulate round them, and are carried with them in their Journeys round their Orbits. These are called Satellites, or Moons, and are ten in Number: of which one well known, belongs to our Earth. Four attend Jupiter; and five wait upon Saturn, who has also a bright Ring surrounding his Body edge-wise, but no where touches it. These Moons (which are a Secondary kind of Planets) are too small to be seen by an Astronomer at the Sun, or even at the Earth when it is much nearer them—without a Telescope. The Tracts too, which they describe in the Heavens, are very complicated; some forming Curves with Loops in them; and others Lines difficult for a Learner to conceive immediately; for which Reason I have reserved the Theory of their Revolutions and Method of calculating their Positions, with respect to their Primaries, for a future Part of this Work.

+ Because no Planets are eclipsed, or hid by one another from the Sun, but when they are in that Plane, and exactly in the same Line continued to the Stars.

The several Periods of the Planets round the Sun, if accurately observed, and reduced to our Time here, as they depart from any fixed Star and return to it again, will be found nearly as under.

	•	Days		Hours	3	Min.	
Mercury	•	87	•	23		16)	,
Venus		224	20	16	-	49	which is the
Earth	revolves about the	365		6	<u>(44),</u>	9 (Length of
Mars	Sun in the Space of	686	etrij.	23	-	27	the Year at
Jupiter		. 4332	120	12	Total	20	that Planet:
Saturn	j . (10759	¥50	6	, 1242;	36 j	•

Mercury and Venus are called inferior Planets, because they are lower or nearer the Sun than the Earth is; but Mars, Jupiter, and Saturn are called superior Planets, because they move beyond the Earth, and are higher in the System.

The Planets he will observe to revolve the same Way, and nearly round the middle or equatorial Parts of the Sun, with a very small Inclination to one another. For if he makes the Earth's Orbit the Standard of the other Orbits, he will observe that they have one Half a little above, and the other Half a little below it, cutting it in two opposite Parts exactly.—The Inclinations of these Orbits to the Plane of the Earth's Orbit he will find to be nearly thus.

The Points where the several Orbits of the Planets cut or cross the Plane of the Earth's Orbit, either in ascending above it, or descending below it, are called the Nodes of those Planets.—That Point where the Planetascends above the Orbit of the Earth is called the North or Ascending Node; and the other Point where the Planet descends below it, is called the Southern or Descending Node.—These Nodes or apparent Intersections of the Orbits of the Planets with that of the Earth's Orb, (if referred to the Ecliptic at the Stars) will be found at this Time to be nearly in the following Places.

The South or Descending Node is in the Sign, Degree, and Minute opposite the ascending one.

THE HEAVENS SURVEYED.

Whilst the Observer views the several Planets revolving round the Sun in their annual Courses, he will perceive that each has a Rotation upon its own Axis, to cause Day and Night at that Planet.—The Times of their several Rotations are as follow.

Mercury
Venus
Earth
Mars
Jupiter
Saturn

D. H. Min.

O - 6 - 0

o - 23 - 0

which is the
Length of the

Day to the Inhatants there.

The Earth turns upon her Axis in respect to the fixed Stars in 23 Hours 56 Minutes, but in respect to the Sun in 24 Hours *.

The Axis of every Planet he will observe to lean or incline to the Plane of its own Orbit, from whence will necessarily arise a Variety of Seasons in those Planets. The Axes of the Planets always keep in the same Position 7 or Parallel round their Orbits. Their Inclinations are as here set down.

The Velocities of the Planets, or Spaces of their Orbs passed thro' in a given Time, he will discover to be greater in the Planets nearer the Sun, and less in those further off. Their Motions in an Hour are as here expressed.

TOURS WILLIAM TO THE TOURS WILLIAM TO THE LICENTERS OF THE TOTAL COLUMN TO THE TOTAL C	THE STATE OF THE S	moves in one Hour about	45.000	and so far are the Inhabit tants of that Planet carried every <i>Hour</i> , without being sensible of that rapid Motion
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From a late Observation, if it may be depended on, the Velocities of the Planets are one sixth Part greater still.

^{*} The Rotations of Mercury and Saturn upon their Axes were observed by the Astronomer Rheita; and Bianchini makes the Rotation of Venus to be 24^d 8h, and the Inclination of her Axis 75°.

† See my Modern Geography or Modern Astronomy, where the Earth, with the proper Inclination of its Axis, is delineated in tour various Parts of its Orbit.

The spectator, whilst he continues at the Sun, will further observe that the Planets in their Revolutions, when they arrive to a certain Part of their Orbits will appear a little larger, and when they come to the opposite Part will appear a little smaller. He will also perceive that when they look smaller their Motion is something slower; but as they begin to appear bigger their Motion grows a little swifter. Hence he will be n turally led to infer, that they must sometimes be a little nearer to him, and at other Times a little further from him; and, consequently, that their Orbits cannot be perfectly circular, but a little eliptical, or oval.—When any Planet is at its greatest Distance, it is said to be in Aphelion (or from the Sun); when at its least Distance, it is said to be in Peribelion (or near the Sun)*.—The exact Places in their Crbits where the several Planets are in their Aphelions, or greatest Distances from the Sun, the observer will find, at this Time, to be when they appear against, or in the following Signs and Degrees of the Ecliptic.

		•	Deg	5 •	Min.
Mercury's	(Sagitarius	13	eli.	45
Venu 's		Aquarius	7	,	37
Earth's	<u> </u>	Capricorn	8	žío	57
Mars's	in or about	Virgo	I	-	55
Jupiter's		Libra	IQ	-	58
Saturn's	n. =	. Capricorn	Ò	(************************************	6

Their Peribelion Points are in the opposite Sign, Degree, and Minute, exactly.

Now, if the Observer be supposed to advance from the Sun to each Planet when in Aphelion, and again when in Perihelion, and accurately measure their several Distances from the Sun to those Points of their Orbits, he will find them to be nearly as under †.

Mercury's Venu's Earth's Mars's	Aphelion, or greatest Dis-tance from	38.455.700 59.418.770 82.368.900 134.421.000	Perihelion, or least Dis- tance from	58.544.300 58.581.230 79.631.100
Jupiter's Saturn's	the Sun	444.290.500	the Sun	403.709.500

^{*} At the Sun, the Planets appear but small: for Mercury's diameter does not subtend an Angle of more than 20"; nor Venus's of more than 30" The Earth appears nearly the same as Venus Mars's Dameter appears about 5" Jupiter's about 40"; and Saturn's about 20". The Satellites or Moons of Jupiter and Saturn, are not visible there.

+ An Observer of the Planets there, might gain some Idea of their relative Distances and Magnix tudes, but could not obtain their real Ones; because no Method offers itself at that Station of ascer-

taining their Parallax: Nay, the Planets seen from thence have no Parallax at all.

When the Planets are in the Middle of their Orbs, between the Aphelion and Perihelion Points,—they are said to be at a mean Distance from the Sun: which Distances of all the Planets will be found nearly these following.

Mercury's Venus's Earth's Mars's Jupiter's Saturn's	Mean or Middle Distance from the Sun is about	32 59 123 424 777	Millions of Miles.
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The Distance of the nearest of the fixed Stars he will find to be almost 20 Millions of Millions of Miles from the Sun, which is more than 20.000 Times farther than Saturn*.

If the Diameters of the Planets be taken by the Surveyor as he passes along in measuring the Distances, they will be found to be nearly as under.

The Diameters multiplied by 3.141 will give their Circumferences.

At the different Planets he will observe the Sun of different Magnitudes; bigger to the nearest, and less to those farther off. The Solar Light is also much greater to the nearer Planets than to those situate higher in the System.

At the fixed Stars the Solar Disk dwindles into a lucid Point, and seems only a Star in the opposite Part of Heaven; and all our Planets are totally invisible there.

The fixed Stars, in all Probability, are so many Suns, having Planets surrounding them, as our Sun has.

As the young Astronomer traverses the System, he will observe that all the Planets cast a Shadow behind them directly opposite the Sun, which Shadow terminates in a Point before it reaches the Orbit of its surrounding Planet: But the Moon revolving round our Earth, and those of Jupiter and Saturn, he will find to run sometimes into the Shadows of their Planets, and by that means be deprived of the Sun's Light. At that Time they become obscure, and are said to suffer an Eclipse.

The Nature of Eclipses of the Earth and Moon are fully explained in the System of Modern Astronomy.

* These are the Distances of the Planets usually received by Astronomers; but if the late Transits of Venus over the Sun may be depended on, the Distances will be found to be about one sixth Part greater; that is, those will exceed these about as much as Statute Miles exceed Computed Miles: But Projections of the System made from either Measure will be equally exact. The Method of calculating the Distances both Ways may be seen in my Trigonometry.

The Sun he will discover to be a huge Sphere, of most wonderful and exquisite Construction, which, by repeated Circumgyrations on its Axis, projects from its Surface sine subtile Particles of Matter (usually called Æther) to the most distant Parts of the Universe.—These Particles by their amazing Velocity constitute not only our Light, but are perhaps the Primum Mobile,—the Spring of Motion and Action of the whole System. Their Motion is inconceivably rapid, being not less than 10 Millions of Miles per Minute; so that they reach Mercury in 3', Venus in 5', the Earth in 8', Mars in 12', Jupiter in 42', Saturn in 77';—but do not arrive at the nearest fixed Star in less than 3 or 4 Years, and to Stars of the Jecond Magnitude not in less than 6 or 7 Years: And so long is Light descending from those Stars to us.—Hence we may restect how very different the Places of the Sun, Planets, and fixed Stars are in the Heavens, from those they appear in from hence.

The Observer, as he leaves the Regions of the Sun and advances toward the Planets, will behold the whole Space he passes through to be obscure and dusky, unless the aforesaid revolving Bodies, which appear very bright, by restecting the Solar Light back to his Eye.—He will also perceive it grow very cold as he recedes from the Sun, because the Light is so subtile as to pervade and pass thro him with little or no Resistance: But as he enters the Atmosphere of any Planet, and becomes invested with the grosser Parts of it, he will find it begin to grow warm, which Warmth will continue to encrease upon him, till he comes down to the Surface of the Planet, where he will experience the greatest Heat of all.

The Heat and Cold at any Planet he will discover to arise—not so much from the Planet's Nearnessto or Distance from the Sun, as from the Nature of its Soil and the Density of its Atmosphere.—Hence it happens that, by the Rarity of it at Mercury, and the Density of it at Saturn, the Inhabitants of the sormer do not complain of the Intensity of the Heat, nor those at the latter lament the Severity of their Cold; but as regular Temperament is observed and kept up at each Planet as upon the Earth where we live. This makes them all fit Places for Habitation, and as comfortable and agreeable in their Variety of Seasons, of Days, and Years, as ours here.

Besides these Wonders of the Heavens, there are many others not yet come to his Inspection*.—To survey them all would be a Work next to impossible; since they are like the Divine Contriver, infinite perhaps, and, at present, past our finding out. But these, already beheld and discovered, are enough to convince us, not only of the Existence of GOD, but also, of his Wisdom, Power, and Goodness; and to make us break out and confess, in the Language of the Psalmist, that—There is no Work like thy Work—for thou art great and doess wonderful Things!

Here

^{*} As (1) What the Power is, which retains the Planets in their Orbits; and keeps the Inhabitants upon their Surfaces as they rowl round their Axes. (2) What the Causes of those Laws are by which that Power acts through all Parts of the System. (3) Whether the same Number of Revolutions of a Planet does not determine the Age of the Inhabitants there. (4) Whether those that live at Mercury are not 4 Times more active than we at the Earth; and whether our Ideas and Actions here are not 30 Times quicker than their's ar Saturn. (5) Whether a Child born at the same Moment on each Planet,—that at Mercury will not be 4 Years old (of our Time);—that at Venus 1 Year 8 Months;—that at Mars Half a Year;—that at Jupiter 1 Month;—that at Saturn only 12 Days, when that upon our Earth shall have compleated 1 Year. These, with Observations relative to the Dimensions of the Inhabitants of each Planet, may be considered in a future Work.

THE HEAVENS SURVEYED.

Here follows A TABLE, in which the foregoing Proportions and Properties of the Primary Planets are brought to one View.

· · · · · · · · · · · · · · · · · · ·			<u> </u>		and the state of t	
Planets Names. Their Revolutions.	Mercury. 87d 23h 16'	Venus. 224 ^d 10 ^h 49'	Harth. 365d 6h 9'	Mars. 686 ^d 23 ^h 27	Jupiter. 4332 ^d 12 ^h 20'	Saturn 10759d 6h 26
Inclination of their Orbits to Earth's.	6° 54 [√]	30 24	the Standard	10 52	1° 20°	2° 30'
Place of the Af- cending Node, the Descending op- posite.	.8 ,15° 45'	11 14° 34'	老 条	18°9'x	8° 32'	21° 25'
Turn upon their	Od ,6h O'	od 23h.c'	od 23h 56	1 ^d 0 ^h 40′,	0 ^d 9 ^h 56'	29 ^d 1.0 ^h 1
Inclination of their Axes to their Orbits		very little-	23° 29'	very little	very little.	30° c'
Place of the Aphe- lions, the Perihe- lions opposite.	13° 45'	7° 37'	8° 57'	my 55'	10° 58'	o° E*
Greatest Distance from the Sun.	38.455.700	59.418.770	82.368.900	134.421.000	444.290.500	821.307.000
Least Distance from the Sun.	25.544.300	58.581.230	79.631.100	111.579.000	403.709.500	732.693.000
Mean Distance.	32.000.000	59.000.000	81.000.000	123.000.000	424.C00.000	777.000.000
Their Diameters in Miles.	2.460*	7.906	7.964	4.444	81.155	.67.870
Motion in 1 Hour.	100.000	70.000	56.000	4-5-000	24.000	18.000
Proportion of Light and Heat.	7 Times greater.	2 Times greater.	at the Earth	½ as much	27 Times less.	90 Times

The mean Distances of the Planets from the Sun, their Diameters, and bourly Motion in their Orbits, as deduced from the late Transit of Venus over the Solar Disk, are as under; and are about one sixth greater than those usually received.

New Distances.	36.841.468	68.891.486	95.173.000	145.014.148	494.990.970	907.956.130
Diameters.	3.100	9.360	○ 843.760○ 7.970	5.150	94.100	77.990
Motion in 1 Hour.	109.699	80.295	68.243	55.289	29.083	22.101

The Distance of the nearest fixed Star is almost 20 Millions of Millions of Millions of Miles from the Sun; which is above 20.000 Times farther than Saturn. Stars of the 2d Size are almost as far again.

The Observer in his Tour thro' the System having made these Observations, and taken these Dimensions, and noted them down in a Table, as a Surveyor does the Angles and Sides of a Piece of Land in his Field-book, he may proceed to delineate the Survey in the following Manner.

Bianchini supposes the Axis of Venus to incline 75 Degrees; that the revolves on her Axis in 24.

Days 8 Hours; consequently has only 9 Days and # in her Year.

* The Light and Heat at the Sun is about 45.000 Times greater than at the Earth.—His Diameter is 763.000 Miles.—The Inclination of his Axis about 8 Degrees, and his Revolution round it is.

À

DELINEATION of the HEAVENS,

ACCORDING TO

The foregoing SURVEY.

(see the two Plates at the Beginning of the Book)—The smaller Part has several Circles described upon it.—The Center represents the Sun.—The sirst Circle next him, the Orb of Mercury;—the second, Venus;—third, our Earth;—fourth, Mars;—fifth, Jupiter;—sixth, Saturn;—and the next Circle (which is double) beyond Saturn, represents the Ecliptic Line at the fixed Stars, divided into the 12 Signs;—and the Circle bounding it is also divided into 12 Signs, and on this is accounted the Earth's Anomaly.—On the other, or larger Part, is described another double Circle, which is divided into 12 Signs. This Circle is used in reckoning the Anomaly of any of the Planets.—This being premised, the Reader may proceed to the Construction or Delineation as follows.

First, make a Scale of equal Parts, of any Size you please;—then taking off about 850 of those Parts, describe the Ecliptic Line, and divide it into 12 equal Parts, and each Part into 30 more; write therein the Names of the Signs and their Characters, as you see in the smaller Plate.—Then you may proceed to delineate the Orbs of the Planets in the following Manner.—— Look in the Table before-going, for the Aphelion Place (first for Saturn) and you will find it to be at o' of Capricorn.—Find this Place on the Ecliptic, and draw à Line from it (with a Black-lead Pencil) through the Center to the opposite Sign and Degree.—Then look also, in the same Table, for the greatest Distance of Saturn, which take off the same Scale of equal Parts, and set it from the Center towards the Aphelion Point, where make a Dot, and write Aphelion. -Take also the least Distance off the Scale, and set it from the Center towards the Perihelion Point, and there make another Dot, and write Perihelion.—Then finding the Middle, betwen those Distances, describe thereon a Circle, which will express the Orbit of Saturn.—In the Table look the Place of Saturn's Node, and having found it in your Ecliptic, lay a Ruler from the Center to that Part,

Part, and where it intersects (or cuts) the Orb of Saturn make a Dot, and inferibe against it N & for North Node. In the opposite Part of the Orbit make another Dot, and insert there S. & for South Node.——In the same Manner you may proceed to project the Orbs, with the Aphelions, Peribelions, and Nodes of all the primary Planets: which being performed, you must then describe a Circle a little beyond the Ecliptic, and right against the Aphelion Point of the Earth begin to divide it into 12 Signs and Degrees, as the Ecliptic is; and number it as you see done there.—On this Line is always found the Earth's Anomaly.

Next, upon the larger Part, describe another Circle, something wider than that of the Earth's Anomaly, and, beginning at any Part, proceed to divide it in every respect as the Earth's is.—On this Circle the Anomaly of the Planets is always found.

Lastly, take the larger Part, cut it round close to the Circle, and paste it upon a strong circular Board of the same Size: This Board should have three small Feet on the under Side, about two Inches high, to raise it free from the Table it shall be placed on when used.—The smaller Part must in like Manner be cut round, and pasted upon a very thin Piece of Board (that will not warp) which just fits it. Then the Centers of both Parts are to be connected by a large Pin or Wire, in such Manner, that the upper Part may turn easily upon the under.

To the Pin, which stands about one Inch high, are to be fixed three fine Ibreads, with a small Plummet or Ball at the End of each, to keep them strait when extended from the Center; and a small Bead may also be put upon each Thread.—Upon the Top of the Pin may be affixed a large Bead, which will serve as well for Ornament, as represent the Sun in the Center of all the Planets.

Thus is the Projection finished, representing the true Proportion and Symmetry of the Heavens; and becomes, thus fitted up, a curious Astronomical Instrument, by which the Places and Situation of all the Planetary Worlds, with respect to one another, imay be determined at any Time, with great Exactness and little Trouble.

Some of the many Uses of this Instrument immediately follow

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OF THIS

Cælestial PROJECTION or INSTRUMENT.

I. To Rectify the Instrument.

IRST, bring the Aphelion Point of the Planet, whose Place you want to find in the Heavens, on the moveable Part, to the Beginning of the Divisions on the Circle in the fixed Part, and let it remain there.

II. To find the Anomalies of any of the Planets.

Turn to the Tables at the End of the Book, and write out the Anomalies of the Earth and of the Planet, answering to the Year, Month, Day, and Hour proposed. These added up severally will be the mean Anomalies* of each for that Time, remembering in Leap Year, after February, to add the Motion of one Day more.

With these mean Anomalies enter the Tables of their Equations, with the Sign at the Top (if less than six Signs), and the Degree in the first Column on the lest Hand descending:——But if the Anomaly be six Signs or more, look the Sign at the Bottom of the Table, and the Degree on the right Hand ascending; and in the common Angle, or Place of Meeting, is the Equation; which, according to the Title, added to, or subtracted from the mean Anomaly, will give the true Anomaly, or real Distance of the Earth, and Planet from their Aphelions at that Time.

III. To find the Heliocentric Place of any of the Planets.

Lay one of the Threads upon the Anomaly of the Planet found in the outer Circle; and also a Thread to the Anomaly of the Earth found in the fecond Circle, so will these Threads severally cut the Heliocentric Places of the Earth and Planet (i. e. the Places as seen from the Sun the Center of the System) in the Ecliptic; and the Places where the Threads cut their Orbits will be the true Places of those two Bodies, and the exact Situation of them in the Heavens at that Time. And if a small Bead, which should be upon each Thread, be brought to each of these Intersections, they will truly represent the Planets themselves.

^{*} The mean Anomaly of a Planet is its Distance from the Aphelion Point, supposing the Orbin which it revolves to be circular. But its true Anomaly is its Distance according to an Ellipsis, which is the true Figure (more or less) of all their Orbits.

To find the Geocentric Place of any Planet, i. e. its Place in the Ecliptic as seen from the Earth.

Lay the Edge of the Scale, or Index, you projected the Instrument by, from the Earth in its Orb to the Planet in his; then bring the third Thread so as to lie parallel to the Index; remembering to lay it ever from the Center towards the Planet; so will the Thread cut the Ecliptic in the Planet's Geoventric Place, and shew the Sign, Degree, and Minute of the Ecliptic it then possesses as seen from bence *.

V. To find the Heliocentric Latitude of a Planet, i. e. its Latitude as seen from the Sun.

At the same Time as the Bead lies upon the Orbit of the Planet, you may easily discover his Distance from the nearest Node; and by knowing the Planet's greatest Latitude, or Inclination, (which is always when he is in the middle, between the Nodes, and is exprest in the foregoing Table) you may nearly estimate his Latitude at that Time.

If the Bead or Thread lie on the right Hand the N. Node, the Planet has North Latitude; and lies above the Plane of the Earth's Orbit, so much as is his Latitude:—If the Thread lie on the right Hand of the S. Node, the Planet has South Latitude; and is below the Ecliptic, or Plane of the Earth's Orbit.

But to discover the Latitude more exact; count the Planet's Distance from the nearest Node in Degrees; with that Distance enter the Table of Inclinations, &c. at the End of the Book; and right against it, under the Planet, you have the Planet's true Heliocentric Latitude at that Time.

The Orb of the fixed Stars being delineated near to the Orbit of Saturn, whereas it should have been placed (had there been Room) above 20.000 Times further off, the Index when laid from the Earth to the Planet, cannot cut the Ecliptic in the Place it would do it it was removed to its true Distance. But as the Distance of the Stars is so immensely great, the Semidiameter of the Earlb's Orbit dwindles as it were into a Point, consequently, a Line extended from the Sun, parallel to the Index, will fall upon the same Point or Part of the Ecliptic that the Index would do. if extended thither.

VI. To find the Geocentric Latitude of a Planet, i. e. his Latitude as seen from the Earth.

The Latitude of a Planet as seen from the Sun, is very different from that seen from the Earth; because the Earth in moving round her Orbit must be sometimes nearer, sometimes farther off the Planet, than the Sun is: But having sound the Latitude as seen from the Sun (by the former Proposition), you may find his Latitude as seen from the Earth in the following Manner.

Measure with the Index the Planet's Distance (in Millions of Miles) from the Sun and from the Earth; then say by the Rule of Three Inverse,

As the Planet's Distance from the Sun, Is to his Latitude seen from thence; So is his Distance from the Earth, To his Latitude seen at the Earth.

Hence it always follows, that when the Earth is nearer the Planet than the Sun is, the Geocentric Latitude will be greater than the Heliocentric: But when the Earth is farther off than the Sun from the Planet, the Geocentric Latitude will be less than the Heliocentric.

VII. To find when any of the Primary Planets are in Aphelion or Perihelion.

Having collected from the Tables the Anomalies for the Time proposed, if the Sum amount to o Sign, o Deg. &c. the Planet is then in Aphelion, or at its greatest Distance from the Sun, and its Motion is then slowest: But if the Sum be exactly 6 Signs, it is then in Perihelion; is nearest the Sun, and its Motion is now at the fastest.

This is seen immediately upon the Instrument.

VIII. To find when a Planet is in Apogeon or Perigeon.

If the Threads when laid-the one to the Earth's Anomaly, and the other to the Planet's Anomaly, lie in a right Line on different Sides the Sun,—the Planet is then in Apogeon, or at his greatest Distance from the Earth:—But if the Threads lie one upon another, on the same Side the Sun, the Planet is then in Perigeon, or nearest the Earth.

F

IX. To find when a Planet is at its greatest Elongation, or Digression from the Sun, on either Side of him.

In the Superior Planets, Saturn, Jupiter, and Mars, this is always when they are opposite the Sun, or 180 Degrees distant from the Sign and Degree he is then in. — But in the Inferior Planets, Venus and Mercury, it is when the Index being laid from the Earth to the Planet does no where cut the Planet's Orbit, but exactly touches it, like a Tangent Line.

X. To know when a Planet is Direct, Stationary, or Retrograde in the Heavens.

(1.) For the Superior Planets.

When the Index is laid from the Farth's Place in her Orbit, to Saturn, Jupiter, or Mars in their's, if the Index does not cut the Earth's Orbit, but becomes a Tangent Line to it, then the Earth is at its greatest Elongation seen from that Planet, and the Planet, about that Time, seen from the Earth, becomes stationary. And if the Earth be advancing towards the Planet on the same Side with the Planet, it is then stationary to Retrogradation; which Retrogradation will become visible in a few Days, and continue all the Time the Earth is on the same Side with the Planet.——And when the Index becomes a Tangent to the Earth's Orbit on the other Side, the Planet becomes stationary to Direction; which Direction of the Planet will take Place in a few Days, and continue all the Time the Earth is on the opposite Side the Sun.

(2.) For the Inferior Planets.

If the Index when laid from the Earth to either of the Inferior Planets, Venus of Mercury, does not cut their Orbit, but touch it only, then you are certain that the Planet is at its greatest Elongation, and stationary at that Time; and, if it lie on the right Hand the Sun, it is stationary to Direction---if on the left Hand, it is stationary to Retrogradation. But if the Index fall within the Orbit, the Planet is (if beyond the Sun) direct in Motion: But if on the same Side the Sun with the Earth, it is going retrograde in that Appearance.

This Rule would hold exactly true, if the Earth were at rest in her Orbit: But as she advances at the same Time the same Way with the Planet, these Stations and Retrogradations will happen later on the left Hand, and some on the right Hand; the sormer arising from the Difference of their Motions; the latter from the Sum: For which Reason the sollowing will be more exact.—Observe the Angle made at the Sun by the two Threads, and if it be Deg.

That Planet is stationary to Retrogradation, if on the left Hand Side the Earth; or station—

That Planet is stationary to Retrogradation, if on the left Hand Side the Earth; or station—

ary to Direction, if on the right Hand Side.—

If the Angle be less, the Planet is retrograde; if bigger, it is direct in Motion.

XI. To discover when any of the primary Planets shine in the Morning or Evening; that is, when they are Morning and Evening Stars.

Bring the Earth's Place in her Orbit so as to lie towards you; and lay the Index across the Instrument from the Earth through the Sun; then all those Planets on the left Hand Side of the Index are occident,* or eastward of the Sun—shine in the Evening, and set after him.—But all those Planets on the right Hand Side are orient, † or westward of the Sun, and shine in the Morning before the Sun rises.

- If a Planet be not farther from the Sun than 12 or 13 Degrees, either before or after him, it will not be visible in the Morning or Evening, by reason of the Sun's superior Light.
 - XII. To know when either of the Inferior Planets, Venus or Mercury, will transit the Sun, or appear as a Spot upon his Face.

If the Threads, when laid to the Anomalies of the Earth and Venus or Mercury, happen to coincide, or lie upon one another, then the inferior Planet must lie exactly between the Earth and the Sun; and if this happen at, or very near the Node of the Planet, it may be seen (with a common Telescope) to pass over the Sun's Face, making a black Spot upon his Disk. But if the Planet be not within a few Degrees of either Node, it will then pass under or over the Sun, just as it has North or South Latitude at that Time.

This being premised, we will proceed to illustrate all with an Example in each Pranet.

- * A Planet is said to be Occident, when it is more towards the East than the Sun; because it is then seen in the occidental, or western Part of the Horizon after the Sun is set.
- † A Planet is said to be Orient, when it is nearer to the West than the Sun; because it appears then in the oriental, or eastern Part of the Sky in the Morning before the Sun rises.

To Calculate the Place of SATURN.

SATURN is the highest Planet in the System, having the Orbits of all the rest included within his Orb. The Time of his Revolution round the Sun, his Distance, Aphelion, Perihelion, Nodes, &c. are expressed in the preceding Table, and accurately projected on the Instrument.

This Planet, to us on the Earth, appears of a dull Lead-colour, and about the Size of Aldebaran, a Star of the first Magnitude. When viewed through a good Telescope, he appears to have a Ring, edge-ways, surrounding his Body, and five small Satellites, or Moons revolving round him beyond the Ring; all which apparently lie in the Space of two Hands-breadths off his Body.

The Times of the Revolutions of these Satellites, with their apparent Distances from Saturn's Body, as observed through a good Telescope, are as here set down.

-	• · · · · · · · · · · · · · · · · · · ·	D. H. M.			
The First next him	revolves	[1:21:19)	[44]	Semi-
Second	about	2:17:40	at the	64	diameters
Third	Saturn	4:12:25	Distance &	123	of
Fourth	•	15:22:41	of	201	Saturn's
Fifth	j in	79: 7:48		[59 [‡]]	Body.

These Moons turn (like Ours) once round their Axes in every Revolution; and consequently keep the same Face always towards Saturn's Body.

The Diameter of his Ring is about 2\frac{1}{4} of Saturn's Diameter.

When Saturn is in the Middle of Gemini and Sagitarius, the Ring appears through the Glass quite open: But when in Pisces and Virgo the Ring seems shut; appearing as a right Line across his Body. The Plane of the Ring is nearly parallel to the Plane of our Equator.

Every Year and 13 Days he is conjoined to the Sun, and appears to move direct, to become flationary, and to run retrograde to us every Year.

His great Distance prevents our discovering his Revolution round his Axis; so that, that Phænomenon remains, at present, not perfectly defined.

His Place in the Ecliptic, at any Time, may be easily calculated in the following Manner*.

EXAMPLE.

^{*} The Calculations of the Places, Immersions, and Emersions of the Satellites of this Planet, with those of Jupiter, will be exhibited in a Work of itself.

THE HEAVENS SURVEYED. 19

EXAMPLE.

Anno 1766, November 1st Day at Noon, I would know (by the Instrument) the Heliocentric and Geocentric Place of Saturn in Longitude and Latitude: His Distance from the Sun and from the Earth: And also, whether he is direct, stationary, or retrograde, at that Time.

Write out the Anomalies of the Earth and Saturn, and equate them from the Tables thus:

Anomaly	of	\in) .		771.2	
	S.		D.		Mi.	Anomaly of 5. D. M.
1761 -	6	arman.	1		34	1761 - 3 - 5 - 5
5 -	11	Res	29	-	42	5 - 2 - I - 2
Nov	9	-	29		37	Nov. ~ 0 - 10 - 10
ist Day -						1st Day 0 - 0 - 2
Mean Anomaly -			ľ			Mean Anomaly - 5 - 16 - 19
Equation subtract	Q	-	1	-	40	Equation subtract 0 - 1 - 42
True Anomaly -	4		0		12	True Anomaly - 5 - 14 - 37
•	The second second	·	- Caristal			12277 2007 W 10

Having rectified the Instrument, by bringing the Aphelion Point of Saturn's Orbit opposite the Beginning of the outer Circle, letting it rest there;—lay one of the Threads on the Earth's Anomaly 4°. 0°. 12' on the upper Part, and it will cut her true Place in the Ecliptic; i. e. 890\f --- Next, lay another Thread to Saturn's Anomaly 5°. 14°. 37'. on the lower Part, and it cuts the Ecliptic in his Heliocentric Place in m 1404. - The Threads remaining in this Position, lay the graduated Edge of the Index from Saturn to the Earth: Bring the other Thread parallel to it (by Help of a Pair of Dividers), and that Thread will cut the Ecliptic in almost 19° of n, which is his Geocentric Place, or Place, at that Time, seen from the Earth, as the other was seen from the Sun*.

Saturn is now retrograde, not only because the Index cuts or falls considerably within the Earth's Orbit on this Side the Sun; but also, because the Angle made by the Threads at the Sun is less than 66°. — His Distance from the Sun, measured with the Index (or a Pair of Dividers), is about 739 Millions of Miles: His Distance from the Earth, about 670 Millions: His Distance from the nearest Node is 3603, which gives, by the Tables at the End, for his Heliocentric Lat. oi Inclination 1°. 29'. South Ascending; but his Geocentric Lat. is 1°. 38'. per Rule at Page 15th.—He thines in the Evening; and his Ring, which is now almost open, may be seen, and one or more of his Satellites, with a tolerable Telescope.

^{*} As the Instrument now lies, there is formed a Triangle, in which is exhibited the Method of calculating by astronomical Tables; for there are given the Angle at the Sun, and (the Logarithms of) the two Sides, to find the Angle at the Planet; which, if the Angle at the Sun be less than 6 Signs, must be added to, but it more, subtracted from the Heliocentric Place to give the Geocentric. This holds in the Superior Planets, but in the Inferior the Angle is applied just contrarywise, as the parallel Thread will shew you at all Times. For that Thread ever forms the same Angle at the Sun, as is made (by the Indix) at the Planet; and falls either to the right or left Hand of the Heliocentric Place, as the Angle at the Planet by its Addition or Subtraction would make it.

To Calculate the Place of JUPITER.

INTEXT within the Orbit of Saturn revolves the refulgent Planet Jupiter. He is a glorious Star of the first Magnitude, and greatly resembles Sirius, the great Dog-star, in Splendor, Colour, and Brightness.

Through an ordinary Telescope, he appears to shine with a full Face, and to have several Belts on his Body; in the lower of which is a large Spot, by which his Rotation upon his Axis was discovered, and that his Axis is nearly perpendicular to his Orbit.—Round his Equator, or middle Parts, may be seen, with the same Glass, to revolve four small Satellites, or Moons, all which, when visible, lie nearly in a Right Line, and within a Hands Breadth or two of his Body: But the Position of that Line has respect to the Situation of the Ecliptic at that Time.

—The outermost Satellite passes wide of the Shadow of Jupiter, two Years in every six; but the other three pass through his Shadow in every Revolution, and are then eclipsed, and also eclipse one another.

Besides these Immersions into his Shadow, they become twice invisible to us in each Revolution (except when the Latitude is too great); that is, once, when they are between the Eye and Jupiter; and again, when they are behind his Body.

Jupiter's Moons, or Satellites, revolve round his Body in the Times, and at the Distances here expressed.

The First Second Second him in
$$\begin{cases}
1 - 18 - 36 \\
3 - 3 - 15 \\
7 - 3 - 59 \\
16 - 18 - 30
\end{cases}$$
 at the Distance of
$$\begin{cases}
5\frac{3}{4} \\
8\frac{1}{4} \\
14\frac{1}{4} \\
25
\end{cases}$$
 Semi-diameters of Jupi-ter's Body.

When Jupiter is our Morning Star, and rises before the Sun, the Spectator sees only the Immersions, or Entrance of the Satellites into Jupiter's Shadow. But when he is our Evening Star, and sets after the Sun, we then see the Emersions out of his Shadow.

Every Year and 33 Days he is conjoined with the Sun, and then, for about 34 Days, lies hid under the Sun's Beams, and consequently becomes, at that Time, invisible to the naked Eye.

The other Phanomena of this Planet are exhibited in the former Table; and his Place at all Times may be easily found by the Instrument, as follows.

To

EXAMPLE.

Anno 1766, November 1st Day at Noon, let it be required to find the Heliocentric and Geocentric Place of Jupiter, his Latitude, and Distance from the Sun and from the Earth, and if he be direct in Motion, or stationary, or retrograde.

OPERATION.

From the Tables collect the Anomalies of the Earth and Jupiter, with their Equations, as under.

Anoma	ly	of	,e).			Anon	aly	of	24	•		
Anoma	# v	S.	•	D.		M.	•		S.	,	D.		M.
1761	444	6	-	1		34	1761	-	4	***	27	•	20
5	-	II	-	29	***	42	5	***	5	46	I	-	42
Nov.	-	9	-	29	-	37	Nov.	-	0	_	25	**	16
ist Day.							ist Day	, .	O	-	0	# * ,	5
Mean Anomaly							Mean Anomaly	• ••	10	,	24	-	23
Equation subtract		.0		1	_	40	Equation add	•	O		3	***	3
True Anomaly	् से '	4	inius	0		1,2	True Anomaly	·	10	کا بروستاندنی خاری	27		26
	-	بد وبسيدات	ين يعجين	التالة سابيا هذه		-			بويدات والأرادا		بالمسلب والمساوعات	-	

Having brought the Aphelion of Jupiter to the Beginning of the outer Circle, lay a Thread to his Anomaly, 10°. 27°. 26′. on that Circle, and it will cut the Ecliptic in \$\pi 80\frac{1}{4}\$, which is his Heliocentric Place.—Another Thread brought to the Earth's Anomaly 4°. 0°. 12′. on the upper Circle, cuts her Heliocentric Place (as before) in \$90\frac{1}{4}*. — Then the Index being laid from Jupiter in his Orb to the Earth in her's, and the other Thread brought parallel to it, that Thread will cut the Planet's Geocentric Place in almost \$\pi 160\frac{2}{4}\$.

Threads is more than $52^{\circ}\frac{1}{2}$.—He is now advancing towards his Aphelion, and shines gloriously in the Morning.—His Distance from the nearest Node is $59^{\circ}\frac{1}{2}$; which gives for his Heliocentric Lat. $1^{\circ}.9'$.—His Distance from the nearest Node is $59^{\circ}\frac{1}{2}$; which gives for Millions of Miles; and from the Earth, about 489 Millions; which makes his Geocentric Lat. $1^{\circ}.2'$, North Ascending.

Jupiter being now a Morning Star, may be seen, and all his Satellites, with a tolerable Glass, and their Immersions into his Shadow (should it so happen at that Time) as they pass behind him.

^{*} The Place of the Sun is always exactly opposite to the Place of the Earth (which the Reader will observe once for all). Consequently, the Heliocentric Place of the Earth, and the Geocentric Place of the Sun, must be continually in Signs and Degrees of the Ecliptic directly opposite to one another. For this Reason it is, that Astronomers speak of the Motion of the Sun; and in their Computations, from their Tables, use the Quantities of his apparent Motion, as if it was real.

To Calculate the Place of MARS.

MARS moves round the Sun in an Orb between the Earth and Jupiter. He appears of a fiery red Colour, and to the naked Eye of different Sizes; sometimes as large as a Star of the first Magnitude; at other Times not scarcely so large as one of the second.

Through the Telescope he appears to increase and decrease in Light like the Moon; and when three Signs distant, is nearly bissected, or divided in two.

This Planet is five Times nearer us at one Time than another; consequently, his Appearance, with respect to his Size, must be so much larger in the former Situation than the latter.—He makes all the Aspects with the Sun, and in Opposition to that Luminary is almost as near the Earth as Venus: Much nearer than Mercury or any of the other Planets *.—Between each Conjunction with the Sun he spends about two Years and 50 Days—and becomes direct, stationary, and retrograde to us, as all the other Planets do.

He does not appear to have any Satellites, or Moons, surrounding him, tho' it is very probable he has one or more,—A broad clouded Belt has been observed to shadow nearly half his Disk; and several Spots have been seen upon his Body, by which his Revolution on his Axis was discovered.

His Distance from the Sun, Period, Aphelion, Node, &c. &c. are inserted in the foregoing Table, and therefore it becomes needless to repeat them here.—His Place in the Heavens, at any Time, is easily found from the Instrument thus.

^{*} As Mars, when in Opposition to the Sun, appears about five Times larger in Diameter than when he is in Conjunction with him; he must be therefore five Times nearer us in one Position than the other: For the apparent Magnitudes of distant Objects increase or decrease in Proportion to their Distances from us; but this Planet Mars keeps always, nearly, the same Distance from the Sun; it is therefore plain, even to Demonstration, that it is not the Earth, but the Sun, that is the Center of his Orbit, and of the System.

EXAMPLE.

Anno 1766, November 1st Day at Noon, what is the Place of Mars in Longinude and Latitude; his Distance from the Sun and Earth; and his other Aspects at that Time, as seen from hence.

OPERATION.

Collect the Anomalies of the Earth and Mars, with their Equations, and set them down thus.

Anoma	ly	of	6	·	,	, ,	Anomaly	of	♂.	·. :
		S.		D.	,	M.		S.	D.	M.
1761	'	6	- COLUM	I	-	34	1761 -	5		
5	1	11	-	29	-	42	5 -	7	- 26	- 5 I
Nov.	106	9	gande	29	-	37	Nov: -	•		
ist Day	#9 '	0	763	0	100	59	1st Day -	Ō	- 0	3I
Mean Anomaly		4	aga.	I		52	Mean Anomaly -	7	~ 2	~ 37
Equation subtract		0	95 0	1		40	Equation add -			
True Anomaly	~ -	4	****	0	er. 10	12	True Anomaly -	7	8	57

Bring the Aphelion of Mars to the Beginning of the outer Circle, and lay a Thread to his Anomaly, as expressed in the Operation above, and it falls in the Ecliptic upon his Longitude or Heliocentric Place, in almost viio. A Thread laid also upon the Earth's Anomaly, gives for her Heliocentric Place 89° 1.

Next, lay the Index from Mars to the Earth, and bring the other Thread parallel thereto; it then cuts his Geocentric Place in = 29½, and there he appears as beheld from the Earth.

The Distance of Mars from the nearest Node is 37° and his Heliocentric Latitude 1°.7'. His Distance from the Sun 114 Millions of Miles, and from the Earth 57 Millions; consequently his Geocentric Lat. is 2°. 14'. South Descending.

This Planet is now an Evening Star, and direct in Motion: And, as he has lately passed his Perihelian, and is on the same Side in Perigeon, or near the Earth, he must appear large and brilliant, like a Star of the first Magnitude. But as the Earth advances surther in her Orb, Mars will apparently grow less; and when they are come on opposite Sides the Sun, his Diameter will be almost five Times less; consequently, his Light will be diminished to one twenty-fifth Part of what it is at this Time.

OBSERVATIONS

TI AVING in the foregoing Part of this Treatise given the Reader a View of the System from the Sun, the Center of it; we will now review it with him from the Earth, the Planet on which he, at present, resides: By this means he will be enabled to account for the various and seemingly irregular Appearances of the Planets; as why they sometimes seem bigger, and sometimes less; why they sometimes move forwards, and backwards, and sometimes seem not to move at all.—In Order to this it will be necessary to delineate the several Planetary Orbs, with the Planets upon them; and then, by transferring the Eye from the Earth to those Bodies, he will instantly perceive all the Phænomena just as in the System itself*.

From the same Scheme it is evident, that to an Eye at the Earth the Planets Mercury and Venus, when on this Side the Sun, must pass between us and him; and with a Telescope may sometimes be seen to make a black Spot upon his Face. But when on the other Side, they pass behind him; and in that Position he may sometimes intervene and eclipse them +.

The dotted Lines drawn from the Earth, touching the Orbs of Mercury and Venus on the Right Hand and on the Left, evidently show that these Points are the greatest Distance which those Planets can ever be seen from the Sun. The Distance of Mercury is never more than 28 Degrees; nor of Venus more than 48. These Distances are called their greatest Elongations on either Side.

Apparent also it is from the Projection, that all the time those two Planets pass from their greatest Elongations on the Right Hand, or Western Side, round the Sun, to their greatest Elongations on the Left Hand, or Hastern Side, they must seem to move through the Heavens according to the Order of the Signs; i.e. from w to &, from & to n, &c. But from their greatest Elongation on the Left Hand, they appear to move the contrary Way, i.e. from n to 8, from 8 to m, &c.... During this State they are said to be retrograde.

When these Planets lie exactly between the Earth and the Sun they are said to be in their inferior Conjunctions; and appear wholly dark, because their upper Sides

* The Moon, which is our Earth's Satellite, revolves round the Earth in 27d. 7h. 43'. Her various Phænomena, with the Moons of Saturn and Jupiter, will be considered in a suture Work.

[†] From an Inspection of this Scheme it is plain, that as a Spectator at the Sun would see the Earth (or any other Planet) revolve thro' the Signs of the Ecliptic; so to a Spectator on the Earth (or Planet) the Sun will apparently revolve the same Way, but always in the opposite Point. For it is well known, that a fixed Object (as the Sun is) appears to change its Place by the Motion of the Observer round it.

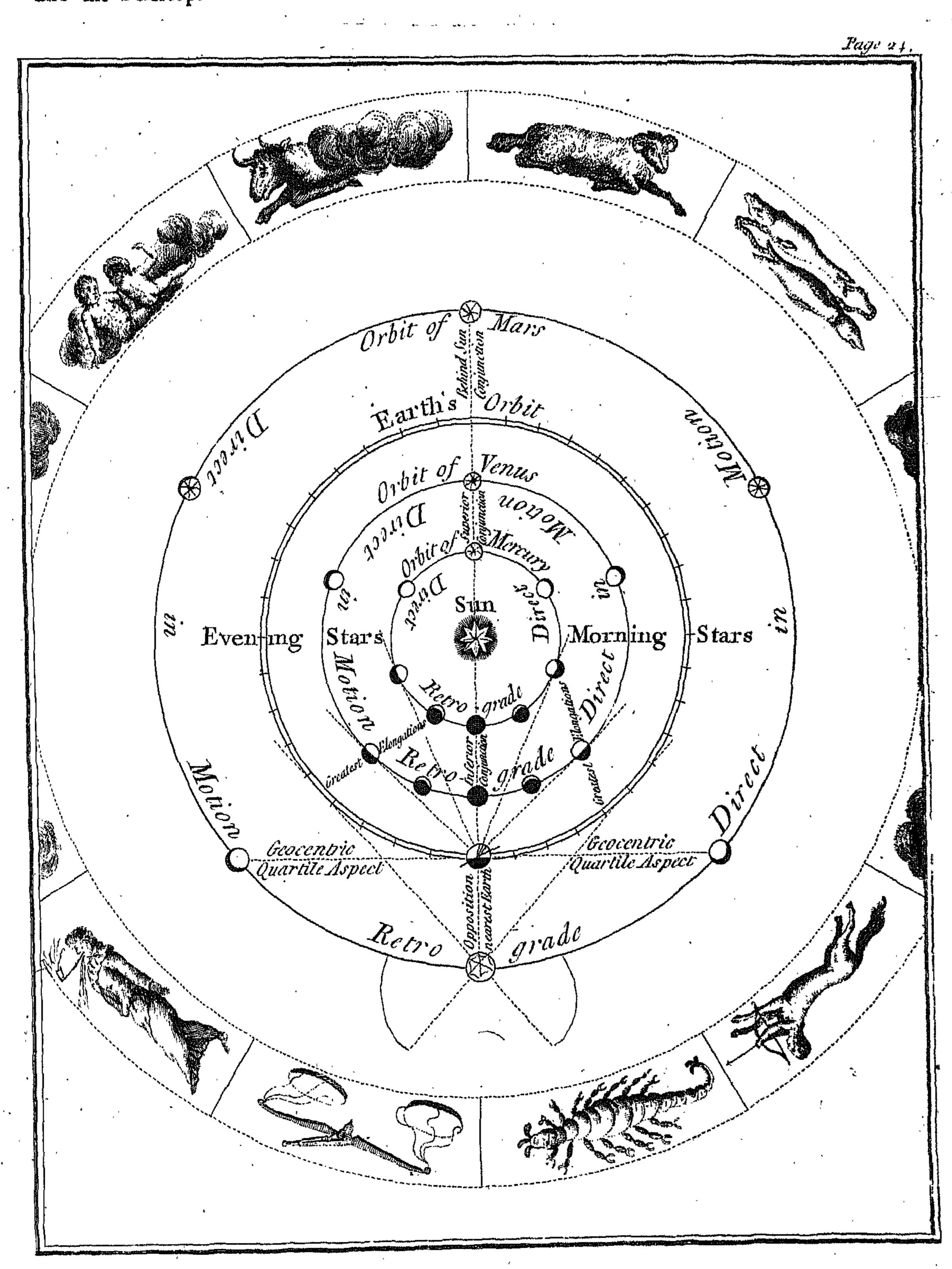
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SYSTEM from the EARTH,

The Place of our present Residence;

In which is clearly seen the true Reason of the Directions, Stations, Retrogradations, and Elongations of the Planets, and the different Phases they exhibit to us (in the different Parts of their Orbits) thro' the Telescope.



only are illuminated by the Sun. But as they recede from that Position, they turn some Part of their enlightened Surface towards us, and appear through the Telescope like a New Moon of three or sour Days old. This Light will continue to increase as they pass round the Sun to their superior Conjunctions, where they turn their whole enlightened Hemisphere towards us. From thence they begin to decrease in their Light, appearing first gibbous, then bissetted, and lastly borned, as delineated in the Scheme.

Though Venus and Mercury (but particularly Venus) a little before their Elongations on the Right Hand, and a little after their Elongations on the Left Hand, turn but a small Portion of their enlightened Surface towards us; yet they will appear more large, brilliant, and glorious, than in any other Parts of their Orbits; because that small Part will apparently contain, at those Places, a greater Area of Light, than the whole Disk when they are at their greatest Distances, four or five Times farther off, beyond the Sun.

These are the principal Phænomena of the inferior Planets Mercury and Venus: Let us proceed to view the superior, Mars, Jupiter, and Saturn. But as the Phænomena of one of them, viz. of Mars, will fully elucidate and explain the the Appearances of them all, I shall, to prevent the Projection from swelling too wide, consider the Revolution and Appearances of that Planet only.

When a fuperior Planet, as Mars, is in that Part of his Orbit nearest the Earth, on the same Side with her, he is said to be in Opposition to the Sun; at which Time all the illuminated Surface of that Planet is turned towards the Earth, and he appears, through the Glass, round like a Full Moon. But as he proceeds towards his Quartile Aspect he decreases in his Light, and appears there gibbous, or a little like a Moon a few Days before the full. As he goes on from thence, he recovers his Light, and shines with a full Face, till he arrives at his Conjunction with the Sun, where he becomes invisible, on Account of the Sun's superior Light. As he comes round from the Sun to his Quartile Aspect, he appears gibbous again, but on the contrary Side; and then recovering his Light as he advances, shines with a full Orb to his Opposition.

This Appearance cannot take Place in Saturn and Jupiter, on Account of their vast Distances, but is very conspicuous in this Planet.

The dotted Lines from Mars to the Earth's Orb, shew the Places where the Earth appears at that Planet to be her greatest Elongation, which never exceeds 34 Degrees.—The Earth, to an Eye at Mars will seem to pass retrograde thro' the inferior Part of her Orb; but direct in Motion through the superior Part; and will appear born'd, bissected, gibbous, and full, through the Glass, just as Venus does to us here. Mars will appear enlightened in like Manner (if visible there) to the Inhabitants of Jupiter; and Jupiter to those of Saturn.

These Appearances, which utterly confound and overthrow the Ptolemaic Hypothesis, are conclusive Proofs of this Copernican System, and cannot fail of establishing it upon a Foundation which shall be as lasting as that of the System itself.

To

To Calculate the Place of VENUS.

NEXT within the Orbit of the Earth revolves the Planet Venus. She is the most splendid and glittering of all the Planets; and so large as to be seen even in the Day Time.— As the Orbit of this Planet is contained within the Orbit of our Earth, she can never come in Opposition to the Sun, (as the superior Planets Mars, Jupiter, and Saturn do) but always continues near him; being never more than 47 Degrees on either Side from him; which Distance is called her greatest Elongation.

When she is orient, i. e. on the Western Side the Sun, she comes to a Conjunction with him by direct Motion in the upper Part of her Orbit; and is then in Apogee, or most distant from us: But when she is occident, i. e. on the Eastern Side the Sun, she comes round to a Conjunction with the Sun by a retrograde Motion, and is then in Perigee, and about six Times nearer the Earth than in the former Situation; that is, nearer by the whole Diameter of her Orbit, as is plain by an Inspection of the Instrument.

Through the Telescope she appears to increase and decrease like the Moon; and is sometime borned and sometimes gibbous; which is a manifest Proof that her Orbit is wholly included within ours.

This Planet, when in Conjunction with the Sun in the upper Part of her Orbit, and near her Node, will pass behind the Sun's Body, and be eclipsed:—But when she is in Conjunction with the Sun in the lower Part of her Orbit, i.e. in the Part next us, and very near her Node, she may be seen, with the Telescope, to pass over the Sun's Face; making a black Patch upon him, like a Beauty-spot in a Lady's Face.

Venus can never be seen at Midnight; but only in the Morning and Evening, just as she happens to be Westward or Eastward of the Sun. When she appears in the Morning, she is called the Morning-star; and when in the Evening, she is called the Evening-star*.—To us here, she appears to be stationary and retrograde once every two Years. Her Period, Distance, Aphelion, and Nodes, and other Affections, are exhibited in the foregoing Table, and her Place in the Heavens is easily obtained in the following Manner.

^{*} This Planet appears brightest and largest at about 40 Degrees from her inserior Conjunction, before and after it.

EXAMPLE.

Anno 1766, November 1st Day at Noon, the Place of Venus is required both Heliocentric and Geocentric; also her Distance from the Sun and Earth; and whether she be direct, stationary, or retrograde at that Time.

OPERATION.

Write down the Anomalies and Equations of the Earth and Venus from the Tables, as follow.

Anomal	y of	€	3 .			Anon	naly	of	Š			
	S,	,	D.		M.	•,		_				M.
1761 -	\cdot ϵ	—	I	*****	34	1761	***	1	673	26	***	22
5 -	· II	•	29	-	42		-					
Nov	, 9	-	29		37	Nov.	· -	4	***	7	-	3
1st Day	- C) -	0	-	59	1st Day		0	-	1	344	36
Mean Anomaly -	- 4	* 	I	ender Time desirent	52	Mean Anomaly		7				
Equation subtract	•				-	Equation add		-				
True Anomaly -	4		O	1	12	True Anomaly	-	7		2 I	بران خزار _{می} ام	7

The Instrument being rectified, by bringing the Aphelion of Venus to the Beginning of the lower Circle; and the Threads laid to their respective Anomalies, will give in the Ecliptic for the Longitude, or Heliocentric Place of Venus $128\frac{1}{2}$, and for the Earth, $128\frac{1}{2}$.—The Index laid from Venus to the Earth, and the other Thread placed parallel to it, gives Venus's Geocentric Place, $122\frac{1}{2}$.

At this Time Venus is distant from the nearest Node about 76°. which gives by the Tables, her Inclination, or Heliocentric Latitude, 3°. 7'. Her Distance from the Sun is nearly 59 Millions of Miles, and from the Earth 130 Millions; therefore her Latitude from hence must be 1°. 29'. North Descending.

Venus is now direct in Motion, advancing towards her superior Conjunction, and will be seen in the Marning before the Sun; at which Time, if viewed with a good Glass, she will appear gibbous, or like a Moon a sew Days before the Full.—She may also be seen in the Day-time by the same Glass, if it be directed properly to her Place in the Heavens. But the best Time to view this Planet by Day is, when she is a sew Degrees before or after her inferior Conjunction; for then she appears borned, like a New Moon, with the illuminated Part turned towards the Sun.

To Calculate the Place of MERCURY.

WITHIN the Orbit of Venus, and next the Sun, revolves the Planet Mercury. He is the nearest to the Fountain of Light of all the primary Planets. He is but seldom seen, because his Orb lies so near the Sun; for at his greatest Distance on either Side him, he is never found further than 27 or 28 Degrees.

This Planet, as well as Venus, can never come into Opposition with the Sun; nor be seen — only in the Evening just after Sun-set, or in the Morning just before he rises.

Mercury becomes direct, stationary, and retrograde three or four Times every Year to us at the Earth; but, like Venus, makes no Aspect with the Sun but his Conjunction.—When he is conjoined with the Sun in the upper Part of his Orb, he is always direct in Motion; and, if near his Node, will pass behind the Sun's Body: But when conjoin'd with the Sun in the lower Part of his Orb, he is ever retrograde; and, if near his Node, he may be seen, with a good Telescope to make a small black Patch on the Sun's Disk.

At his inferior Conjunction, Mercury is nearer the Earth than at his superior Conjunction by the Diameter of his Orbit.—He appears the smallest of all the primary Planets;—is of a sparkling red Colour, greatly resembling the Planet Mars, but not quite so large.—His lying so near the Sun, makes it impossible for us to make the same Observations on him with the Telescope as on the rest of the Planets: For we have not been able (on Account of the Heavens being so strongly illuminated by the Sun when he is visible) to discover from hence any Spots on his Body, or if he revolve on his Axis or not.

The Time of his Revolution round the Sun, Distance, Places of his Nodes, Aphelion, &c. are before expressed in the general Table: and his Place, at any Time, is easily calculated as follows.

EXAMPLE.

Anno 1766, November 1st Day at Noon, suppose it be required to find by the Instrument the Place of Mercury in Longitude and Latitude, as seen from the Sun and from the Earth; also his Distance from those Places; and whether he be direct, stationary, or retrograde, &c.

OPERATION.

From the Tables collect the Anomalies of the Earth and Mercury, and their Equations, as here set down.

Anomaly	of	€	} •			Anon	aaly	of	å	•		
_						-		S.		D.		M.
1761 	· 6	-Q+	I		34	1761	gap	8		2	52 0	43
5			•			5	pm.	9	•	2	27	36
Nov.	7		7		- A- A-	Nov.	<u>po</u>	5	hw	14	•••	4
ist Day -	Q	P#	Q	(Files	59	1st Day	•	Q	**	4	*	6
Mean Anomaly -						Mean Anomaly	e.jp	10	Zipr Zipr	23	igas galanes cill	29
Equation subtract	· Q	444	I	;;==	40	Equation add	;**	0	S 42	11	S	32
True Anomaly -	4		O		12	True Anomaly	, Alter	I I		5		

Bring the Aphelion of Mercury (as you did the rest of the Planets) to the Beginning of the outer Circle; lay the Threads to their Anomalies severally, and they will cut the Ecliptic in 891° . for the Place of the Earth, and $m18^\circ$, for the Place of Mercury, as seen from the Sun.—Lay the Index from the Earth to Mercury (which is almost a Right-line); bring the other Thread parallel to it, and it will fall upon $m121^\circ$. his true Place seen from the Earth.

Mercury is now about 3 Degrees from the Node. His Inclination, or Latitude at the Sun, is 22'. His Distance from the Sun is 36 Millions of Miles, and from the Earth 116 Millions; which gives his Latitude, as seen from the Earth, about 7'. South Descending.

He is now in the upper Part of his Orbit; is in Apogeon; has just past his superior Conjunction, and is direct in Motion; but lies so near the Sun as not to be visible, rising and setting nearly at the same Time with him.

If two white Beads, having that Half blacked opposite the Sun, be brought, one of them to the Earth's Place, the other to the Planet's in their Orbs, you may pleasantly see how much of that Planet appears to be illuminated at the Earth, and also what Aspect the Earth appears in at that Planet.

To Calculate a TRANSIT of VENUS over the Sun's Disk.

As the Orbit of Venus is wholly included in the Earth's Orbit, she must sometimes interpose between us and the Sun; and if that Interposition should happen at either of her Nodes, (that is, just at the Intersection of her Orbit with the Plane of the Earth's Orbit) she must apparently pass over the Sun's Face.—The Places where these two Intersections are made are at 14 Deg. 34 Min. of Gemini, and 14 Deg. 34 Min. of Sagitarius.—The former is the Place of her Ascending Node; i. e. where she ascends above the Plane of the the Earth's Orbit;—The other is her Descending Node, or Place where she descends below it. Transits of this Planet can therefore only happen in those Months in which the Earth is in these two Signs, and near the Degrees where the Nodes are situate; that is, in the Beginning of June and December: At all other Seasons of the Year, that Planet will pass under or over the Sun's Disk, without producing that agreeable Phænomenon.

Now it has been found by Astronomical Calculation, that Venus will return to the same Situation with the Earth every 8 Years, 235 Years, and 243 Years nearly: Transits will therefore happen at the End of these Intervals; but more exactly as expressed in the following Tables.

A TABLE of the Periods of VENUS's Retrograde Conjunction with the Sun at the Ascending Node in the Month of December.

```
Years. D. H. M. Latitude of Venus. 

8—subt. 2: 10: 52^{\frac{1}{2}} - - 24': 41" South. 

235—add 2: 10: 9 - - 11: 33 North. 

243—subt. 0: 0: 43 if Common Year 

243—add 0: 23: 17 if Leap Year } 13: 8 South.
```

Angle of the visible Way of Venus over the Sun's Disk

Semidiameter of the Sun at that Time - - - - 0:16:21

The Hourly Motion of Venus within the Sun - 0:4:7

Consequently, the greatest Duration of the Transit is - 7^h:56':0"

Venus was within the Sun's Disk at this Node, New Style, 1639, December 5 D. 6 H. 37 M. with 8 M. 30 Sec. South Latitude.

A TABLE of the Periods of VENUS's Retrograde Conjunction with the Sun at the Descending Node in the Month of June.

Angle of the visible Way of Venus over the Sun's Disk - 8°: 28': 0"

Semidiameter of the Sun at that Time - 0: 15: 50

The Hourly Motion of Venus

And, the greatest Duration of the Transit is - 7^h: 76': 0"

Venus was seen within the Disk of the Sun at this Node, 1761, June 5 D. 17 H. 55 M. with 4 M. 15 Sec. South Latitude.

^{*} For the Rationale of these Numbers, see Dr. Halley's Discourse in the Philosophical Transactions, No. 193,—or Whiston's Lectures.

From the Periods of Venus here laid down, it will be very easy to compute the Time of any suture Transit, by Addition and Subtraction only. For if to the Time of any Transit given, you add or subtract the first Period, and increase or decrease the Latitude, as much as the Table directs at that Period, you will have the Return of the next Conjunction of Venus with the Sun at that Node, with her Latitude at that Time; which if less than the Sun's Semidiameter, there will happen another Transit.

If the first Period does not bring the Latitude within the Limits prescribed, you must try with the second Period; and if that does not do, apply the third Period: Always remembering to use the first before the second, and the second before the third; as in the following Example:

EXAMPLEI.

I would know the Time of the next Transit of Venus over the Sun at the Ascending Node in the Month of December.

O P E R A T I O N.

Year. D. H. M. M. Sec.

Venus in the oat Ascending 2 1639 Dec. 5: 6: 37 Lat. 8: 30 South.

First Period - add 8 subt. 2: 10: 52* 24: 41 South.

1647 Dec. 2: 19: 45 33: 13 South.

Hence it appears that Venus will pass below the Sun's Disk, because her Latitude exceeds the Sun's Semidiameter, 16'. 21". by many Minutes: We must therefore try the second Period.

Year. D. H. M. M. Sec.

Venus in the oat Ascending 8 1639 Dec. 5: 6: 37 Lat. 8: 30 South.

Second Period - add 235 add 2: 10: 9 + 11: 33 North.

1874 Dec. 7: 16: 46 3: 3 North.

At this Time Venus will be seen in the Sun, and will pass over his Disk, about 3 Minutes northward of his Center.

EXAMPLE II.

I would also know when the next Transit of Venus will happen at the Descending Node in the Month of June.

OPERATION.

Venus in the \odot at Descending \odot 1761 June 5:17:55 Lat. 4:15 South. First Period \odot add 0:10:58 North

1769 June 3: 10: 57 15: 43 North.

These are all the Transits of Venus that will happen in this and the next Century.

^{*} In this Example you add the Years, but subtract the Days, Hours, and Minutes.

† Here you must subtract the less Latitude from the greater, because they are different: But when alike (i. e. both North, or both South), add them together.

From

To Calculate a TRANSIT of MERCURY over the Sun's Disk.

of the Earth, there must necessarily happen the like Appearances in this Planet as in that; but they are more frequent in Mercury than in Venus, because of his quicker Circulation, and speedier Arrival to the same Places with the Earth. The Nodes of this Planet, i. e. the Places where his Orbit cuts the Plane of the Earth's Orb are in 15°.45', of Taurus, where he ascends above the Ecliptic; and in 15°.45' of Scorpio, where he descends below it. All Transits of Mercury must happen in the Beginning of May and November, when the Earth is in those Signs where his Nodes are.

The several Periods of Mercury's Conjunction with the Sun, both at the Ascending and Descending Node (as determined by Astronomical Observations and Calculations), are exhibited in the following Tables.

A TABLE of the Periods of MERCURY's Retrograde Conjunction with the Sun at the Ascending Node in the Month of November.

```
Years. D. H. M.

6—add 8: 17: 25 If 2d or 3d past Bissextile
6—add 9: 17: 25 If Bissextile, or 1st past
7—subt. 7: 0 add 9 If 2 Bissextiles in the Time
7—subt. 6: 0 add 9 If 1 Bissextile - 22': 47" South.

13—add 2: 17: 34 If 3d after Bissextile - 3" North.

The Angle of Mercury's visible Way over the Sun - 8": 15': 0"
The Sun's apparent Semidiameter at that Time - 0: 16: 5
```

Mercury was in the Sun at this Node, New Style, 1756, Nov. 6 D. 16 H. 36 M. with 0 M. 39 Sec. South Latitude.

A TABLE of the Periods of MERCURY's Retrograde Conjunction with the Sun at the Descending Node in the Month of May.

The Angle of the visible Way of Mercury over the Sun - 10°: 18': 0"

The apparent Semidiameter of the Sun - 0: 15: 46

Mercury was in the Sun at this Node, New Style, 1753, May 5 D. 19 H. 20 M. with 1 M. 19 Sec. South Latitude.

In the same Manner as you calculated the Times of the future Transits of Venus, you may calculate, by these Tables, those of Mercury: An Example or two will make all plain.

EXAMFLEI.

Suppose you would know the next Transit of Mercury over the Sun at the Ascending Node in the Month of November.

As Mercury (at the Time of the last Transit) had such small Latitude, I perceive immediately, from the Table, that the Latitude of Mercury at the two first Periods is too great to produce that Phænomenon: For at the End of the first Period of six Years he will go North, and in the second of seven Years South of the Sun's Disk: I therefore make use of the third Period thus.

	Years.				-		M. Sec.
Mercury was in the o	1756	Nov.	6;	19	: 36	with Lat.	o: 39 S.
The third Period -	13	add	2:	17	34		8: 3 N.
Next Transit at Ascending 8				•	10	· •	7:24 N.
,	nagara tanàna bandara	كبينتك بيمشنمه لدكسته	سبقته ومبدؤق	ملطلب بدولتها بالم	entirella mindle		Angelist seems - construction of the party problem.

E X A M P L E

When will Mercury transit the Sun next, at the Descending Node in the Month of May?

In this Example, the Latitude of Mercury carrying him above or below the Disk in the two first Periods, I apply the third, as under.

		D. H.		M. Sec.
Mercury in the Oat Descending a 3d Period, because Leap Year,	3 1753 Ma	y 5:19;	20 with L	at, 1:19 S.
3d Period, because Leap Year, a	idd 33 fub	t 2; 0:	23	•
Next Transit at Descending &	1786 Ma	y 3; 18;	57	12:43 N.

In this Example I subtracted one Day more on Account of its being Leap Year.

When the Latitudes are both of one Name, add them together; but when of different Names subtract the smaller from the larger, and the Remainder will be Latitude at that Time.

The other Transits of Mercury which fall within this Century, are,

Years.	D.	H.	М.		Sec.	
1776 Nov.	2;	10	1.9	15	: 23	South Latitude.
1782 Nov.	ĮI:	3	44	15	: 27	North Latitude.
1786 May	3 :	18	57		•	North Latitude.
1789 Nov.	-			7	: 20	South Latitude.
1799 May	G.			4	: 12	South Latitude,

If to any of these Periods of Venus or Mercury you find the Planet's Anomaly with the Earth's, and lay the Threads thereto, you will find them lie upon each other; which proves them to be, at that Time, in Conjunction, if viewed from the Sun: Consequently, the Planet seen from the Earth (as it is then near the Node) will appear in the Sun's Disk.

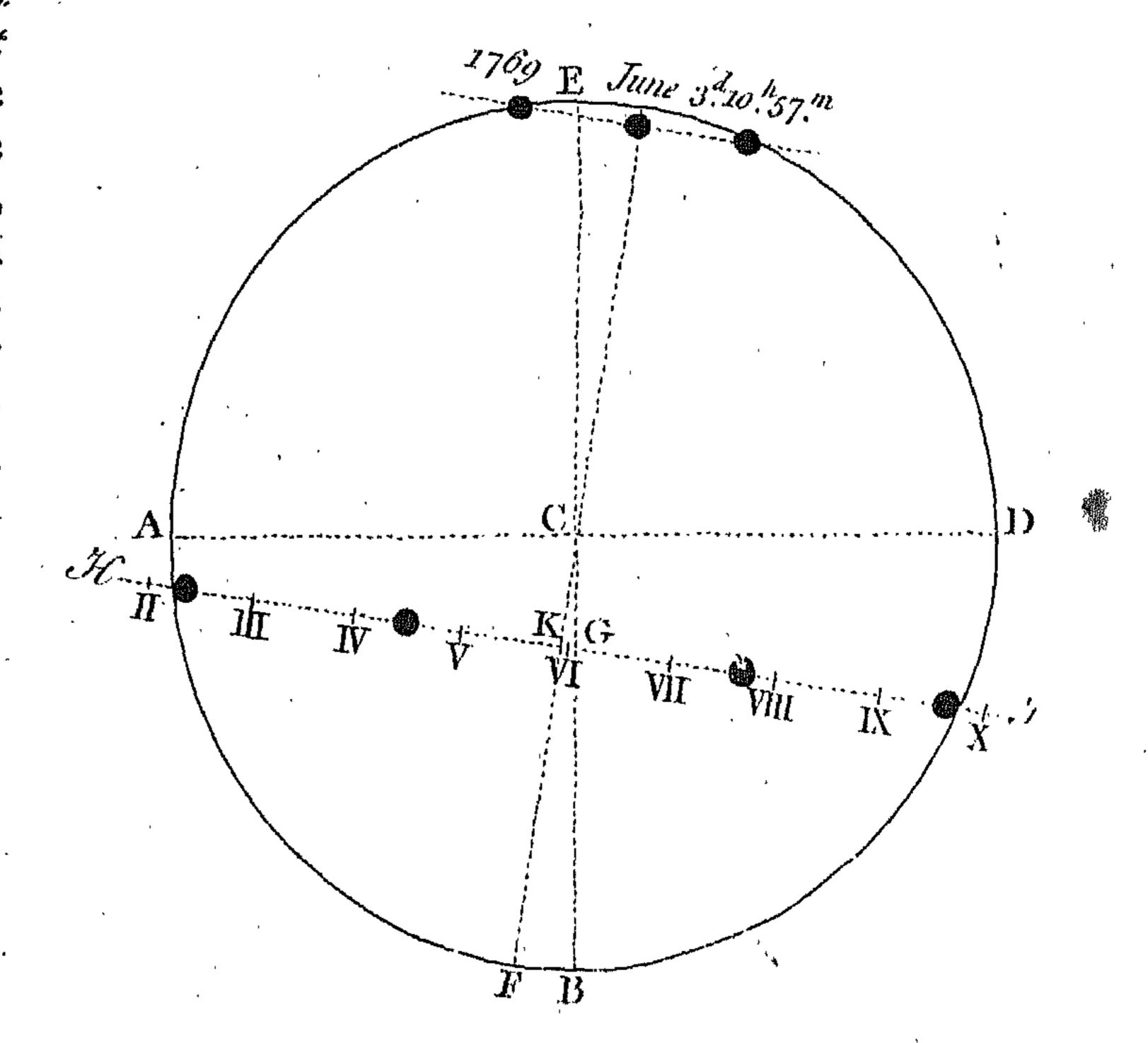
To Delineate a TRANSIT of VENUS (or Mercury) over the Sun's Disk.

(Suppose that which happened June 5th, 17 H. 55 M. or the 6th Day, 55 M. past 5, in the Morning, 1761, with 4'. 15". South Latitude).

TROM any Scale of equal Parts, take off, with a Pair of Dividers, 15'.50".

= the Semidiameter of the Sun, and draw the Circle ABDE, which will represent the Sun's visible Disk.—Through the Center C draw the Line AD, to represent the Ecliptic, and cross it in the Middle with the other EB, for the Axis of the Ecliptic.—Then with a Protractor, or Line of Chords, set off the Angle of 8°.28'. = the Angle of the visible Way of Venus from B towards F, because that Planet is retrograde, and Latitude Ascending, or Increasing South; and draw CF for the Axis of Venus's Way.—Take the Latitude 4'.15". at the Time of that Conjunction, from the same Scale of equal Parts, and set it on the Axis of the Ecliptic from C to G, and draw HI through the Point G, at Right Angles to CF; so will HI describe the visible Path of Venus over the Sun.

And as the longest Time of any Transit of Venus is when the passes over the Middle of the Sun at AD, which is performed in 7.56', or eight Hours nearly, you have, by dividing the Diameter into eight equal Parts, the Space that Venus déscribes upon the Disk in one Hour. And as the Middle of this Transit, at K, happens at 55' past 5, -- take off 5' from one of the Hours, which you have previously divided into 60 equal Parts, and set, them



from K towards G; that will show where Venus will be upon the Disk at 6 o'Clock.—Next, take off the Length of one whole Hour, and set it from 6, each Way along the Path HI. By this Means you will obtain the Hours both before and after 6; which number'd as in the Scheme, will shew the Time of the Beginning of the Transit at H, and the End at I.—You may draw little Circles upon the Path, to represent Venus upon the Disk. Her Diameter is about i'15", which is equal to one of the Divisions of the Scale of equal Parts by which you projected the Disk of the Sun.

The Path at the upper Part of the San represents the Transit of Venus which happened 1769, June 3 D. 10 H. 57 M. with 15'. 43". North Lat. Descending.

These are all the Transits of this Planet within this Century.

** These Calculations and Projections are according to the mean Motions and mean Latitudes of those Planets, and may differ some little from the true Time; but will serve to illustrate and explain to the young Astronomer the Nature and Manner of a Transit, and enable him better to understand the Calculations from the exactest Astronomical Tables.

In all Transits at the Descending Node the Planets pass over the Sun descending, and at the Ascending Node they pass the Disk ascending.

To

To find the Times of the OPPOSITIONS and CONJUNCTIONS of the Primary PLANETS.

SUPPOSE the Earth in Conjunction, &c. with any of the Superior Planets, Mars, Jupiter, or Saturn; then because the Earth revolves with a quicker Motion in her Orbit than either of those Planets, she will every Day recede from them, so much as is the Difference of their Motions; and when she has performed her Period, she will not come up with them again, because they have advanced with their Motion at the same Time: The Earth therefore must go on still further to overtake those Planets, and come to a Conjunction again: And to find how much, say by the Rule of Proportion—

As the Difference between the Earth's and Planet's Motion for any Time, suppose for one Day:—Is to that Time::—So are 360°, the whole Circle:—To the Time they meet again.

By comparing the daily Motions of the Earth and the other Planets inserted in the foregoing Table at Page the 10th, and making the above Proportions; we find the Time between two Conjunctions of the Earth and Saturn to be 378 Days, or 1 Year and 13 Days—Between two Conjunctions of the Earth and Jupiter, 398 Days, or 1 Year and 33 Days.—Between two Conjunctions of the Earth and Mars, 2 Years and 50 Days.—Between two of the Earth and Venus of the same Kind*, 1 Year and 218 Days.—And between two of the Earth and Mercury of the same Kind, 115 Days.

There is the same Time between two Conjunctions of the *Planets* with the Sun; or between any other two similar Aspects with Him.—And the Time between the Opposition and Conjunction of a Planet with the Sun is equal to half the Time of their Conjunctions specified above.

The Times of these Conjunctions added to the former Conjunction found in any Almanack, &c. will give the Time of the next Conjunction following.—But as these Periods are computed according to their mean Motions, i. e. supposing them to move always in circular Orbits, (whereas they move in elliptical ones, moving sometimes a little faster, and sometimes a little slower;) it may happen that the true Conjunctions may fall a Day or two sooner or later than by the Calculation here given: But the true Time will be always shewn by the Instrument.

^{*} By two Conjunctions of the same Kind in Venus and Mercury is meant two Conjunctions either in the Superior, or two Conjunctions in the inferior Parts of their Orbits.——Every Conjunction of Venus is made about 9 Signs further in the Ecliptic, and Mercury about 4 Signs; but the Time in Mercury is a little uncertain on Account of his great Eccentricity.

Of the Direct, Stationary, and Retrograde Appearances of the Primary PLANETS.

THE Stations and Retrogradations of the Planets in the Heavens are not real, but apparent only, being caus'd by the Revolution of the Earth, and the various Positions of it in its Orbit, as was shewn before in treating of that Planet; and therefore it becomes unnecessary to repeat it here.

Tho' it is difficult to determine the exact Times of these Appearances by Calculation, yet the Instrument will shew them at all Times.—For if the Index when laid from the Planet's Place in its Orbit to the Earth touches the Earth's Orb, that Planet appears stationary in the Heavens at that Time. But if the Index cuts the Earth's Orbit, and the Earth and Planet are on opposite Sides the Sun, the Planet is then direct; but if both of them are on the same Side, the Planet is then retrograde.

To estimate nearly the Times of the Directions, Stations, and Retrogradations, &c. of the Planets, I have here subjoined the following Table, which the young Astronomer will find of much Use to him, in understanding and ascertaining these Appearances.

The Planets.	From one con- junction, Sta- tion, and Re- trogradation to another.	Days Direct.	Days Retro- grade.	Degrees Re- trograde.
Saturn Jupiter Mars Venus Mercury	1': 13' 1: 33 2: 50 1: 218 0: 115	238 279 700 542 92	140 120 80 42 24	7 10 20 15 10* 18 16 8;

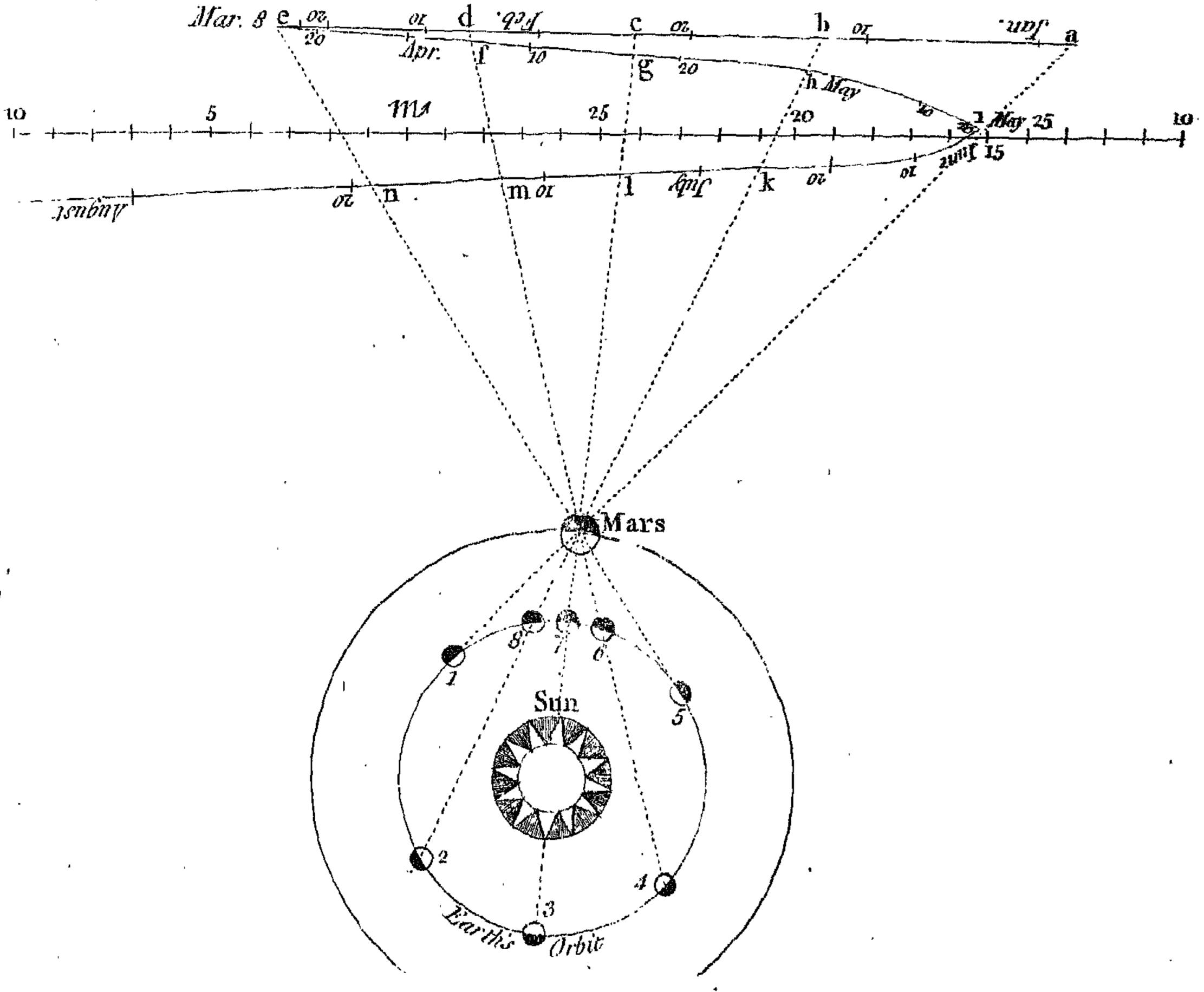
This Table is so clear and evident that it needs no Explanation: For you see that Saturn after 1 Year and 13 Days returns to the same Aspect with the Earth (or Sun);—continues direct in Motion 238 Days;—then runs retrograde 140 Days;—and goes back 7 Degrees in the Ecliptic +.

A Planet in his retrograde Motion does not describle exactly the Tract back again which he passed through in his direct Motion, because the Latitude is continually altering as seen from the Earth; so that sometimes he may pass a little above the Ecliptic Line; sometimes a little below it.—To explain this more fully, I shall here subjoin the direct and retrograde Course of the Planet

^{*} Mars and Mercury are variable in their Retrogradations on Account of the great Eccentricity of their Orbits.

[†] Fromondus, who was averse to the annual Motion of the Earth, was obliged to confess that, that Motion was maintained by the Copernicans by no Argument more probable and specious than that of the Station, Direction, and Retrogradation of the Planets.—Whiston's Lectures, p. 203.

Mars, which he took in the Year 1762, and also the Place where it was stationary; by which the Learner will be enabled to understand those Appearances the better.



In this Scheme, the Center represents the Sun; the inner Circle the Orbit of the Earth, the other the Orbit of Mars.—The Earth being at 1, Mars is seen among the fixed Stars at a.— When the Earth is at 2, Mars is seen at b. When the Earth is come to 3, 4, and 5, Mars appears at c, d, and e.—Thus from a to e Mars moves direct, i. e. from West to East in the Heavens.—But the Earth being at 5, Mars is seen at e to be stationary, or at a Stand, for a few Days.—As the Earth moves from 5 to 6, Mars is seen to return in a retrograde Direction from e to f.—The Earth going on from 6, to 7, 8, and 1, Mars still proceeds to go retrograde from f, to g, h, and i.—At i Mars appears stationary again for a few Days.—The Earth removing from 1, to 2, 3, 4, and 5, Mars is seen to recover his direct Motion again, and to remove from i, to k, l, m, and n.

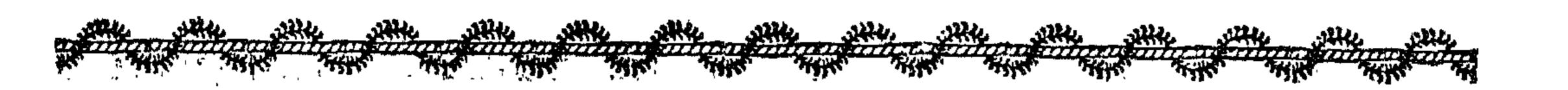
I have supposed Mars in this Projection to remain fixed in the same Place of his Orbit, in order to render the direct and retrograde Appearances evident to the Pupil; tho' it is certain he advances in his Orb the same Way, and at the same Time; but it would be very difficult to express in a Drawing that compound Motion. But the direct and retrograde Lines at the Top, to which the Months are affixed, is the apparent Tract which that Planet described in 1762. and the dotted Line is a Portion of the Zodiac it passed thro' in that Time.— At the End of the preceding Year Mars was observed at a, in about 14° of Libra, with 2° of North Lat. from whence he proceeded direct thro' the Months January, February, and to the 8th of March; when he became stationary in about 3° 3' of Scorpio, with 2° 32' North Lat.—Thro' the Remainder of March, April, and to the 25th of May, he proceeded retrograde, his Latitude decreasing, at which Time he became stationary again in about 150 of Libra.— From hence he began to move direct again, cross'd the Ecliptic Line in about 15° 10' of Libra, and went on thro' June, July, August, &c. with South Lat. to the End of the Year.

To determine the greatest Elongations of the EARTH from the SUN, as beheld from any of the Superior Planets; and also, the greatest Elongations of the Inferiors from the SUN as seen from the EARTH.

BRING a Thread from the Sun across the Instrument, so as to cut the Orbits of all the Planets;—then lay the Index from the Intersections of the Thread with the Orbits of the superior Planets to the outer Part of the Earth's Orbit; and the several Angles measured at the Planets with a Protractor, (or by bringing another Thread parallel to the Index) will point out the greatest Elongation of the Earth to each of these Planets:—And which you will find to be at Saturn about 6 Degrees:—At Jupiter about 11 Degrees:—At Mars about 42 Degrees.—But for Venus and Mercury, lay the Index from the Intersection of the Thread with the Earth's Orbit to the Edge of their Orbits, and measure in like Manner the Angle at the Earth, and you'll have for the Elongation of Venus about 46 Degrees;—and Mercury about 22 Degrees*.

As these Elongations of Venus and Mercury are the greatest Distances these wo Planets ever appear from the Sun, either before or after him; so the above Elongations of the Earth from the Sun is the greatest Distance that ever she appears to those Planets from him.—At Mars the Earth will never be seen to go farther from the Sun than about 42 Degrees, which is not so far as Venus is seen to do to us.—At Jupiter the Distance of the Earth will never be more than 11 Degrees from the Sun, which is but half the Distance that Mercury is found to be at here from the Sun; consequently a Sight of the Earth, even to them, must be a very rare and unusual Phænomenon.—And at Saturn the Earth can never be found to be more than 6 Degrees from the Sun; which Distance is not above one fourth of what we observe in Mercury: And as Mercury is seldom seen here, it is very probable that the Earth can never be seen there; and that the Saturnian Astronomers (if any) have not discovered that there is such a Planet or World as our Earth in the System.

^{*} Mercury, on Account of the great Eccentricity of his Orbit, will have very different Elongations:—Sometimes it will be about 27 Degrees, at other Times not more than 17 Degrees.——But the Elongations in the other Planets will seldom differ more than 1 Degree.



NEW

ASTRONOMICAL TABLES:

SHEWING,

The Mean Anomalies of all the Primary Planets,

WITHTHE

EQUATIONS of their ORBITS;

By which their True Distances from their Aphelions may be easily ascertained at any Time.



NEWTABLES,

Shewing the Mean Anomalies of all the Primary Planets (or their Distance from their Aphelion Points), in Years of Christ current, adapted to the New Style.

**************************************					· · · · · · · · · · · · · · · · · · ·							
	These Numbers correspond to the Old Style.											
Years of Mercury Venus Earth Mars Jupiter Saturn												
Christ.	ğ	\$	Θ	₹	24	ħ						
<u> </u>	80	5 0 /	's , o /	8 0	5 0	5 0						
1 1881	7 19 40	I 29 53	6 13 11	11 19 55	8 0 41	6 18 45						
1701	8 4 11	8 3 25	6 12 59 6 12 47	7 7 5 ² 2 25 49	4 7 34 0 14 28	10 27 6						
172E 1741	0 10 44	8 το 28	6 12 35	10 13 46	8 21 21	7 1 16						
-	Marring N	Jumbers a	aree to t	he Mega	Stule whi	ch com						
1116 10	The following Numbers agree to the New Style, which com- menced September the 2d, 1752.											
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	s o	5 0	5 0	ک دن د مدید سے سو	3.0	5 0						
1761	8 2 43 8 17 14	I 26 22	6 1 34 C	5 25 57 1 13 54	4 27 20 1	3 5 5						
1801	8 27 40		6 0 9		- T	7 13 24						
182x	9 12 10	8 5 20	5 29 57	_ 1	5 1.7 56	3 17 34						
1841	9 26 41	2 8 52				~ ' ~ '						
1861	10 11 12	8 12 21	5 29 33	7 25 14	10 1 44	7 25 45						

A TABLE of the Mean Anomalies of all the Primary Planets in Months.

Months.	Mercury §	Venus P	Earth	Mars 8	Jupiter 24	Saturn
January February March April -	3 0 0 0 0 0 4 6 52 8 1 27 0 8 19	s o o o o o o o o o o o o o o o o o o o	0 0 0 I 0 33 I 28 9 2 28 42	0 0 0 0 16 15 1 0 55 1 17 10	s o o o o o o o o o o o o o o o o o o o	0 0 0 0 1 2 0 1 58
June July August	4 11 5 8 17 57 0 20 43 4 27 34	6 15 15 8 1 55 9 19 59 11 9 39	3 28 16 4 28 50 5 28 24 6 28 57	2 2 53 2 19 8 3 4 51 3 21 6	0 9 58 0 12 33 0 15 2	0 4 I 0 5 3 0 6 3
September - October - November - December -	9 4 26 1 7 12 5 14 4 9 16 50	0 29 19 2 17 23 4 7 3 5 21 7	7 29 30 8 29 4 9 29 37 10 29 11	4 7 20 4 23 4 5 9 18 5 25 2	0 20 11 0 27 41 0 27 43	8 8 0 0 8 0 0 10 10 0 10 11 0 11 0 11 0

In Leap Years, after February, remember to take the Motion of one Day more.

A B L E

Of the Mean Anomalies of all the Primary Planets in Years compleat.

Years, compleat	Mercury	Venus P	Earth	Mars o	Jupiter 4	Saturn
1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 15 15 19 20	10 26 19 0 20 0 2 17 48 4 11 31 6 5 13 7 28 55 9 26 42	1 15 29 9 15 25 9 15 25 7 16 11 3 0 58 10 15 41 6 2 7	II 29 II II 29 55 II 29 40 II 29 24	7 26 51 2 8 7 8 19 23 3 1 11 9 12 27 3 23 43 10 4 59 4 16 46 10 28 2	5 1 42 6 2 2 7 2 21 8 2 45 9 3 24 11 3 46 9 3 24 10 4 28 2 4 47 3 5 6 4 5 50	2 15 14 2 25 26 3 7 40 3 19 52 4 2 4 4 26 30
40 60 80 100	0 29 2 1 13 33 1 28 3 2 12 34	0 7 3 6 10 25 0 14 6 6 17 28	II 29 36 II 29 24 II 29 12 II 29 0	3 5 54 10 23 51 6 11 49 1 29 46	4 13 47 0 20 41 8 27 35 15 4 28	4 8 21 0 12 32 8 15 42 4 20 52

A T A B E

Of the Mean Anomalies of all the Planets in Days.

Days	ğ	Ş	Θ	₹	24	Ď
	3 0	s 0 /	\$ 9	0 /	0	0
Y	0 4 6	0 1 36	0 0 59	0 31	0 5	0 2
2	0 8 11	0 3 12	O I 58	1 3	0 10	0 4
3	0 12 17	0 4.48	0 2 57	1 3	0 15	06
4	0 16 22	0 6 25	0 3 57	2 6	0 20	0 8
5	0 20 28	081	0' 4 56	2 37	0 25	OIO
6	C 24 33	0 9 37	0 5 55	3 9	0 30	0 12
7	0 28 39	0 11 13	0 6 54	3 40	0 35	0 14
8	I 2 44	0 12 49	0 7 53	4 12	0 40	0 16
9	1 6 50	0 14 25	0 8 52	4 43	0 45	0 18
IO	1 10 55	OIO	0 9 51	5 14	0 50	0 20
II	I 15 I	0 17 37	0 10 50	5 46	0 55	0 22
12	1 19 6		0 11 50	6 17	I O	0 24
13	1 23 12	0 20 50	0 12 49	6 49	I 5	0 26
i ¹ 4 i	I 27 I8	,	0 13 48	7 20	1 10	0 28
1 15	2 I 23	0 24 2	0 14 47	7 52	115	0 30
16	2 5 29	0 25 38	0 15 46	8 23	. I 20	0 32
17	2 9 34	0 27 14	0 16 45	8 55	1 25	0 34
18	2 13 40	0 28 50	O 17 44	9 26	1 30	0 36.
19	2 17 45	I 0 26	• • •	9 57	1 35	0 38
5.0	2 21 51	1 2 2	0 19 43	10 29	1 40	0 40
21	2 25 56	1 3 39	0 20 42	i i o	I 45	0 42
22	3 0 2	I 5 15	0 21 41	11 32	_	0 44
23	3 4 7	I 6 51	0 22 40	12 3	I 55	0 46
24	3 8 13	1 8 27	J J/	12 35	2 0	0 48
25		1-10-3	0 24 38	13 6.	2 5	0 50
26	3 16 24	1 11 39	0 25 38	13 37	2 10	0 52
27	3 20 30	I 13 15	0 26 37	14 9	2 15	0 54
28	3 24 35	1 14 52	0 27 36	14 40	2 20	0 56
29	3 28 41	1 16 28	0 28 34	15 12	2-25	0 58
30	4 2 46	1 18 4	0 29 34	15 43	2 30	1 0
31	4 6 52	1 19 40	I 0 33 1	16 15	2 35	I 2
32	4 10 57	I 21 16	1 1 32	16 46	2 40	I 4
	<u> </u>				•	

In Leap Year, after February, remember to add the Motion of a Day more.

A T A B L E

Of the Mean Anomalies of all the Planets in Hours.

	ğ	2	Θ	8	24	5
		• /	0	0 /	0 /	0
I	0 10	0 4	0 2	O I	0 0	0 0
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3	0 31	0 12	0 7	0 4	OI	0 0
4	0 42	0 10	0 10	0 5	OI	0 0
5	5 1	0 20 1	0 12	0 7	OI	0 0 1
. 0	1 1	0 24	0 15	0 8	0 1	0 0
7	IIZ	0 28	0 17	0 9	0 1 1	O I .
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14	2 23	0 56	0 34	0 18	0 3	
15	2 34]	1 0 [0 37	0 20	0 3	0 1
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17	2 54	1 8	0 43	0 22	O' 4	OI
18	3 4	I 12	0 44	0 24	0 4	OI
19	3 14	r 16	0 47	0 25	0 4	0 2
20	3 25	1 20	0 49	0 20	0'4	0.2
21	3 35	1 24	0 52	0 27	04	0 2
22	3 45	1 28	0 54	0 29	0 5	0 2
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The Anomalies of the Planets in Minutes.

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40	0 7	. 0	3	0	2	0	I	0	0	0	o '
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The

The EQUATION

Of MERCURY'S ORBIT.

Mean i	S	ubtracted fro	om the Mea	n Anomaly g	ives the Tr	ue.	Mean f
cury's Anomaly	o Sign	ı Sign	2 Sign	3 Sign	4 Sign	5 Sign	Anomaly
0 1 2 3 4 5 6 7 8 9 0 1 1 2 3 4 1 5 6 7 8 9 0 1 2 2 2 2 2 2 2 2 2 2 2 2 2 3 2 3 3 3 3	D. 0 0 0 0 1 1 1 2 2 2 3 3 3 3 4 4 4 5 5 5 6 6 6 7 7 7 7 8 8 8 8 9 9 9	D. M. 9 53 11 30 10 30 10 30 10 40 11 11 11 50 10 10 10 10 10 10 10 10 10 10 10 10 10	D. M. 17 48 18 18 18 18 18 18 18 18 18 18 18 18 18	D 57 38 38 22 29 2 23 35 7 90 I 2 2 3 3 3 3 4 4 4 2 2 2 2 3 3 3 3 4 4 4 2 2 2 2	D. 48 42 23 4 4 4 31 9 7 4 0 6 1 6 0 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	D. 14 38 26 0 3 36 30 34 44 33 20 9 37 56 5 5 44 3 32 21 10 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	3 9 8 7 6 5 4 3 2 1 0 9 8 7 6 5 4 3 2 1 0 9 8 7 6 5 4 3 2 1 0 9 8 7 6 5 4 3 2 1 0
	II Sign	10 Sign	9' Sign	8 Sign	7 Sign	6 Sign	
() () () () () () () () () ()	Add	to the Mc	an Anomaly	gives the	True.		

The EQUATION

Of VENUS' ORBIT.

Ven Mean A	St	ubtracted fro	om the Mea	n Anomaly g	gives the Tr	ue.	Mean.
Venus' ean Anomaly	o Sign	1 Sign	2 Sign	3 Sign	4 Sign	5 Sign	Anomaly
O 1 2 3 4 5 6 7 8 9 0 1 1 2 3 4 5 6 7 8 9 0 1 1 2 3 4 5 6 7 8 9 0 3 3 4 5 6 7 8 9 0	D. 0 0 1 2 3 4 4 5 6 7 8 9 9 0 0 1 1 2 3 3 4 4 5 6 7 8 9 9 0 0 1 1 2 3 3 4 4 5 6 7 8 9 9 0 0 1 1 2 3 3 4 4 5 6 7 8 9 9 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	D. 23 22 22 22 22 23 23 23 23 23 23 23 23 2	D. M. o 41 0 42 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	D. M. 88 47 77 77 77 77 76 66 66 66 66 67 77 77 77	D. 41 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	D. 24 0 23 0 22 1 0 0 18 8 7 6 0 16 0 0 16 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 0 17 6 Sign	398 76 24 32 10 98 76 54 32 10 98 76 54 32 10

The EQUATION

Of the EARTH'S ORBIT.

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	1 2 3 4 5 6 7 8 9 0 1 1 2 3 4 5 6 7 8 9 0 1 1 2 3 4 5 6 7 8 9 0 1 1 2 3 4 5 6 7 8 9 0 1 2 2 2 2 2 2 2 2 2 2 3 4 5 6 2 3 3 3 3 4 5 6 7 8 9 0 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 5 0 2 3 5 7 8 0 1 1 3 4 6 7 9 0 2 3 5 6 7 8 9 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 39 1 41 1 42 3 44 55 1 45 1 45 1 45 1 45 1 45 1 45 1 4	156 156 156 156 156 155 155 155	1 40 1 38 7 6 1 37 1 38 1 38 1 38 1 38 1 38 1 38 1 38	0 57 5 53 I 98 6 42 0 8 6 42 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	298 76 54 32 1 0 98 76 54 32 1. 198 176 154 132 11 0 98 76 54 32 1.

EQUATION

Of Mars's Orbit.

Mean A	Subtr	atted from	the Mean	Anomaly	gives the	True.	Mean
ars's Anomaly	o Sign	r Sign	2 Sign	3 Sign	4 Sign	5 Sign	Anomaly
0 I 2 3 4 5	D. M. 0 0 0 10 0 19 0 29 0 39 0 49	D. M. 4 50 4 59 5 16 5 25 5 34	D M. 8 41 8 47 8 53 8 58 9 9	D. M. 10 35 10 36 10 37 10 38 10 39 10 39	D. M. 9 44 9 35 9 29 9 24 9 18	D. M. 5 54 5 33 5 23 5 12 5 1	30 29 27 26 25
56 78 9 10 11 12 13	1 1 29 1 39 1 49 1 59 2	5 4 ² 5 5 5 8 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	9 15 9 25 9 35 9 35 9 35 9 44 9 48	10 40 10 39 10 39 10 38 10 38 10 37 10 36	9 9 6 9 6 9 53 40 8 8 33 8 25	4 39 4 39 4 4 5 4 5 3 4 3 30	24 23 22 21 20 19 17
14 15 16 17 18 19 20	2 18 2 28 2 38 2 48 2 58 3 16	6 48 6 56 7 12 7 19 7 27 7 34	9 56 9 56 10 4 10 7 10 10 10 13	10 34 10 31 10 29 10 27 10 24 10 22	8 18 8 10 8 2 7 46 7 38 7 29	3 18 3 54 2 42 2 29 2 17 2 5	16 15 14 13 12 11
21 22 23 24 25 27 28	3 3 4 4 3 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	7 48 7 48 7 55 2 9 8 28 8 28	10 10 10 22 10 24 10 27 10 29 10 31	10 19 10 13 10 9 10 6 10 2 9 58	7 12 7 7 2 7 53 6 44 6 34 6 14	I 52 I 40 I 28 I 15 I 3 O 50 O 37 O 25	98 76 5 4 3 2
29	4 40 4 50	8 35 8 41 10 Sign	10 34 10 35 9 Sign	9 54 9 49 9 44 8 Sign	6 4 5 54 7 Sign	0 12 0 0	I
		Add to the	e Mean A	nomaly give	es the True		

The EQUATION Of JUPITER'S ORBIT.

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niter's	o Sign	r Sign	2 Sign	3 Sign	4 Sign	5 Sign	nomaly
OI2345678 90112345678 9012222456 278 293	D. 0 5 0 6 6 1 2 7 2 3 8 3 8 3 8 4 4 5 9 5 10 5 11 5 1 5 1 5 1 7 2 2 2 2 2 3 3 7 2 3 7 3 7	D. M. 372 47 16 1 50 149 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3	D. 38 1 4 4 7 9 2 5 5 7 0 2 4 7 9 1 3 4 6 8 9 1 2 2 3 5 6 7 8 8 9 9 0 0 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	D 31 31 31 30 0 98 7 7 25 4 32 20 9 7 15 4 20 7 5 3 0 8 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	D. 5520744444444444444444444444444444444444	D 54944 9 33 8 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
	11 Sign	10 Sign	9 Sign	8 Sign	7 Sign	6 Sign	
		Add to the	Mean Ar	nomaly gives	the True.		

The EQUATION

Of SATURN'S ORBIT.

Mean /	Subtracted from the Mean Anomaly gives the True.						
Saturn's Anomaly	o Sign	1 Sign	z Sign	3 Sign	4 Sign	5 Sign	Anomaly
O # 2 3 4 56 78 90 11 2 3 4 56 78 90 1 2 2 2 2 3 4 56 78 90	D. 0 6 12 9 3 18 4 4 5 6 7 3 9 6 12 8 4 4 7 3 9 7 11 7 3 2 9 9 4 7 3 8 4 7 3 8 4 7 3 8 4 7 3 8 4 7 3 8 4 7 3 8 4 7 3 8 4 7 3 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	D. 40 51 6 2 2 3 3 7 2 8 3 8 3 8 3 8 3 8 3 8 3 8 3 8 3 8 3 8	D. M. 270 3471 4470 368 1 468 1 135778 0 2 2 3 46 78 9 9 0 1 1 1 1 1 2 2 2 3 4 6 7 8 9 9 0 1 1 1 1 1 2 2 2 3 4 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	D. M. 31 1 2 2 2 2 3 3 3 3 3 3 3 2 2 2 3 3 3 3	D. 58 44 1 7 40 6 2 8 14 9 5 0 6 1 6 1 6 1 6 1 6 2 1 9 4 8 2 6 0 4 8 8 1 9 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	D. 226 936 92 33 33 33 3 3 3 2 2 2 2 2 2 2 2 2 1 1 1 1	30 98 276 25 4 2 2 2 2 2 2 1 1 1 1 1 1 1 1 1 1 1 1 1
Add to the Mean Anomaly gives the True.							

A T A B E

Shewing the Inclinations, or Heliocentrick Latitudes of all the Primary Planets, by having their Distance from the nearest Node.

nearest in Des	3	Q.	3	21.	
from Node	D M.	D. M.	D. M.	D. M.	D. M.
O	0 0	0 0	0 0	0 0	0 0
I	0 7	0 4	0 2	OE	0 3
2	0 15	0 7	0 4	0 3	0 5
3	0 22	OII	0 0	0 4	0 8
4	0 29	0 14	0 8	0 0	O 10
5	0 36	0 18	U 10	0 7	0 13
6	0 44	0 21	012	08	0.6
7	0 50	0 25	0 14	o IC	0 18
8	0 58	0 28	0 16	OFE	0.21
9	1 5	0 32	0 17	0 13	0 24
10	1 13	0 35	0 19	0 14	0 26
II	I 20	0 39	0 21	O 15	0 29
12	I 27	0 42	0 23	0 16	0 31
13	1 34	0 46	0 25	0:8	0 34
1 1	III	0 49	0 27	(0 19	? 36
1 15	1 48	0 53	0 29	0 21	0 39
16	I 55	0 56	0 31	0 22	0 41
17	2 2	0 59	0 32	0 23	0.44
18	29	I 3	0 34	0 25	0 46
19	2 16	16	0 36	0 26	0 48
20	2 23	1 9	.038	0 27	0 51
21	2 30	1 13	0 40	0 29	0 54
22	2 37	1 16	0 42	0 30	0 56
23	2 43	I 19	0 43	031	0 59
24	2 50	1 23	0 45	0 33	II
2.5	2 57	I 26	0 47	0 34	I 3
2.5 2.6	3 2	I 29	0 49	0 35	1 6
27	3 10	r 32	0 50	0 36	1 8
28	3 16	I 35	0 52	0 38	III
2.9	3 23	I 38	0 54	0 39	1 13
30	3 29	1 42	0 56	0 40	I 15

T A B

Shewing the Inclinations, or Heliocentric Latitudes of all the Primary Planets, by having their Distance from the nearest Node.

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100		•			
from Node grees.					
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30	3 29	I 42	0 56	0 40	1 15
31	3 36	144	0 57	0 41	I 17
32	3 42	1 48	0 59	0 42	1 20
33	3 48	1 51	T O	0 44	1 22
34	3 54	1 53	I 2	0 45	1.24
35	4 0	1 50	I 4	0 46	1 26
36	4 6	I 59	I 5	0 47	1.28
37	4 I 2	2 2	I 7	0 48	1 30
38	4. 18	2 5	I 8	0 49	I 32
39	4 23	2 8	1 10	0 50	I 34
40	4 29	2 I I	TII	0 51	I 37
41	4 35	2 13	I 13	0 52	I 39
42	4 40	2 16	I 14	0 54	I 41
43	4 46	2 18	. I 10	0 55	I 42
AA	4 41	221	I 17	0 50	-I 44
45	4 56	2 24	I 19	0 57	1 46
46	5 I	2 27	I 20	0 58	I 48
47	5 6	2 29	121	0 59	I 50
48	5 11	2 31	1 23	0 59	I 5 ²
49	5 16	2 33	I 24	I O	I 53
50	5 21	2 36	1 25	(A)	1 55
51	5 26	2 38	1 26	I 2	I 57
52	5 30	2 40	I 27	1 3	1 58
53	5 35	2 42	I 29	I 4	2 00
54	5 39	2 44	1 30	r 5	2 I
55	5 43	2 46	1 31	I 5	2 3
50	5 47	2 48	1 32	16	2 5
57	5 5 I	2 50	x 33	I 7	2 6
58	5 55	2 52	I 34	I 8	2 7
59	5 59	2 54	I 35	I 8	2 9
60	6 3	2 56	1 36	I 9	2 10

A T A B E

Shewing the Inclinations, or Heliocentric Latitudes of all the Primary Planets, by having their Distance from the nearest Node.

•				5 	
nearest Nod in Degrees.	ğ	\$	8	4	
from Node	D M.	D. M.	D. M.	D M.	D. M.
60	6 3	2 56	1 36	1 9	2 10
6 r	6 7	2 58	1 37	I 10	211
62	6 10	2 59	1 38	1 10	2 13
63	6 13	3 1	1 39	III	2 14
64	0 17	3 3	1 40	I 12	2 15
65	6 20	3 4	1 41	1 12	2 16
66	6 23	3 0	141	1 13	2 17
67	6 26	3 7	1 42	I 13	2 58
68	6 29	3 0	43	I 14	2 19
09	0 31	3 10	1 44	1 14	2 20
70	0 34	3 11	44	I 15	2 2 1
71	6 39	3 12	45	1 15	2 22
72	6 39 6 4 I×	3 13	I 40	I 10	2 23
73	6 42	3 14	I 40	I 16	2 24
	6 16	3 10	I 47	1 7 7	2 24
75 76	6 47	2 17	1 48	1 18	2 25
77	6 49	2 18	1 48	1 13	2 26 2 26
78	6 50	3 19	r 48	r 18	2 27
79	6 52	3 20	I 40	1 18	2 27
80	6 53	2 20	I 40	1 1X 1	2 28
81	6 54	3 21	1 50	1 18	2 28
82	6 55	3 21	I 50	1 19	2 29
83	6 56	3 22	1 50	1 19	2 29
84	6 57	3 22	r 51	r 19	2 29
85	6 58	3 23	I 51	1 10	2 30
86	6 58	3 23	1 51	I 19	2 30
87	6 59	3 23	I 51	1 19	2 30
88	6 59	3 23	I 51	1 19	2 30
89	6 59	3 23	1 51	1 19	2 30
90	0 59	3 23	I 51	1 19	2 30
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The foregoing Tables are adapted to the New Style*. The first Table, i. e. for the Years of Christ current, exhibits the Distances of the Planets from their Aphelions on the last Day at Noon of the preceding Year. For from thence it is Astronomers begin their Year, (which holds true also of the Months and Days) and not from Midnight of the last Day, as the civil Year begins.—Astronomers likewise reckon on regularly through the twenty-four Hours, from the Noon of one Day to the Noon of the following Day. Thus, January 10th Day 18th Hour in the astronomical Accompt, answers to the 11th Day 6 o'Clock in the Morning of the common or vulgar Reckoning. This should be carefully minded, else you will be liable to commit great Mistakes.

It may be necessary further to observe, that because the Aphelions of the Planets have a slow progressive Motion in the Echptic, but on the Instrument are fixed to one particular Point; hence it will come to pass, that in an Age or two, the Places found by this Method will differ a few Minutes from their true Places in the Heavens †.——But then the

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Instrument

^{*} The Numbers are adapted to the New Style, only to give the Learner less Trouble in reducing the Old Style to the New before he begins his Calculations, as all Tables are accommodated to that Kind of Reckoning.

[†] It has been doubted by many Astronomers, whether the Aphelions and Nodes of the Planets are moveable or not -Street, Whiston, and even Newton, asserted that they are at Rest, with respect to the fixed Stars; and consequently, that they advanced only so much as is the Recession of the Equinox (backwards) Copernicus, I believe, with Halley and Leadbetter, made the Aphelions advance forward much faster, and with a Motion different to each other. The Stars seem to advance only I Degree in 72 Years; but Saturn's Aphelion (according to Leadbetter) proceeds 1 Degree in 45 Years: Jupiter's, 1 Degree in 50 Years: Mars's, 1 Degree in 51½ Years: Venus's, 1 Degree in 63 Years: and Mercury's, i Degree in 69 Years. And upon this Supposition the foregoing Tables of the Planets Distances from their Aphelians are constructed.—But if, at last, it should be proved that the Aphelions do not advance in their Orbs, but continue immoveable with respect to the fixed Stars, then it will be easy to project the Instrument in such Manner as shall make it perpetual; and it will become almost as useful as the large Volumes of Tables, in ascertaining the true Places and Situations of the primary Planets in all Ages from the Creation to the Flood,—to the Time of Hesiod,—to Eudoxus,—to Aratus,—to Hypparchus, to Ptolemy, --- to Copernicus, --- to Kepler, --- to Newton, --- and to the End of the great Platonic Year:

Instrument may be easily projected a new, by the Rules laid down in the former Part of the Book: And, though it should err a few Minutes, it will still have its Uses in pointing out the Situation of the Triangles made by the Lines in the Heavens at the Time of the Calculations, and thence evidently inform the Learner of the true Theory of the System; and by that Means make him Master of this fine Science; which is not so easily to be attained, perhaps, by any other Method.

The Survey of the several Satellites attending the Primary Planets,—as that belonging to our Earth;—the Four of Jupiter's, and the Five of Saturn's; with the the Position of his Ring at all Times, will be exhibited in a suture Work.

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